



Magnetic confinement fusion: science
that's hotter than a Kardashian
Instagram post*

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*Thanks to ChatGPT

Honourable mentions:

- From Plasma to Party: A Fun and Easy Guide to Magnetic Confinement Fusion
- Fusion Energy: A Recipe for Success, Just Add Heat and Magnetism
- The Power of Fusion: Not Just a Dance Move Anymore
- The Fusion Chronicles: A Tale of Science, Magic, and High Temperatures

Most energy on Earth comes from fusion

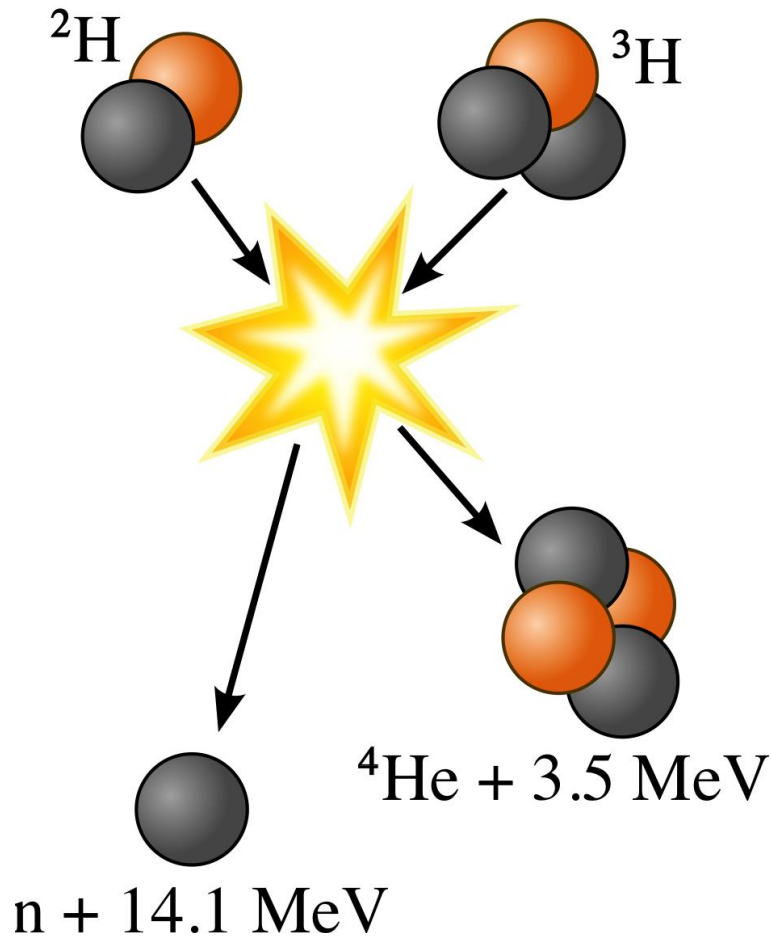


Credit: Diyana Dimitrova/shutterstock.com



Credit: Chris Lim, wikipedia

The fusion reaction

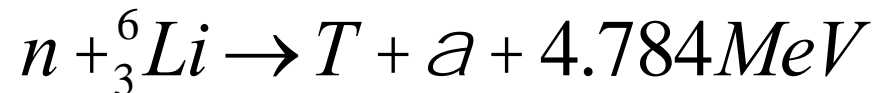


$$E = mc^2$$

$$(m_D + m_T)c^2 = (m_a + m_n)c^2 + 17.6 \text{ MeV}$$

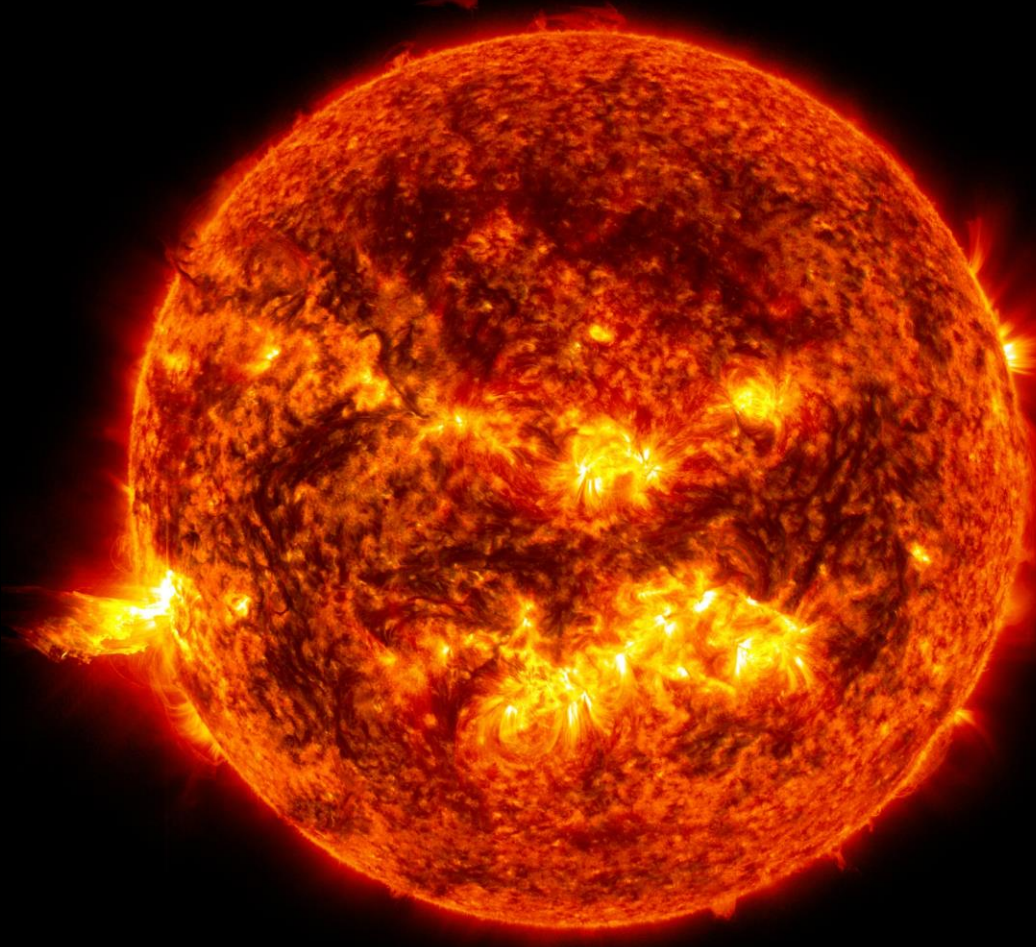
10^7 times energy released
by gas combustion!

Caveat: must breed Tritium



Limited by Lithium supply ~20k yrs

Controlled thermonuclear fusion



- **Must overcome Coulomb barrier to fuse nuclei**
- **Thermonuclear fusion imparts necessary energy via thermal energy**
- **Need about 100 Million degrees Celsius to get significant fusion reaction rate (hotter than Sun)**
- **Must find way to initially heat hydrogenic nuclei and must insulate long enough that fusion is self-sustaining**

Magnetic confinement fusion



© Carol Sexton Art

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1 gram H at 100M° for 1 second

$$\text{Power out} = \frac{nT}{\tau_E} = \text{fusion power in} = \mathcal{S}_{\text{fusion}}(n, T)$$

$$\mathcal{S}_{\text{fusion}} = n^2 \langle \sigma v \rangle E_{\text{fusion}}$$

Fusion cross section peaks at $T \sim 100\text{M}$ degrees

$$\text{Macroscopic stability} \quad \longrightarrow \quad n \sim 10^{20} / \text{m}^3$$

$$\text{Microscopic stability} \quad \longrightarrow \quad V \sim 1 \text{ m}^3$$

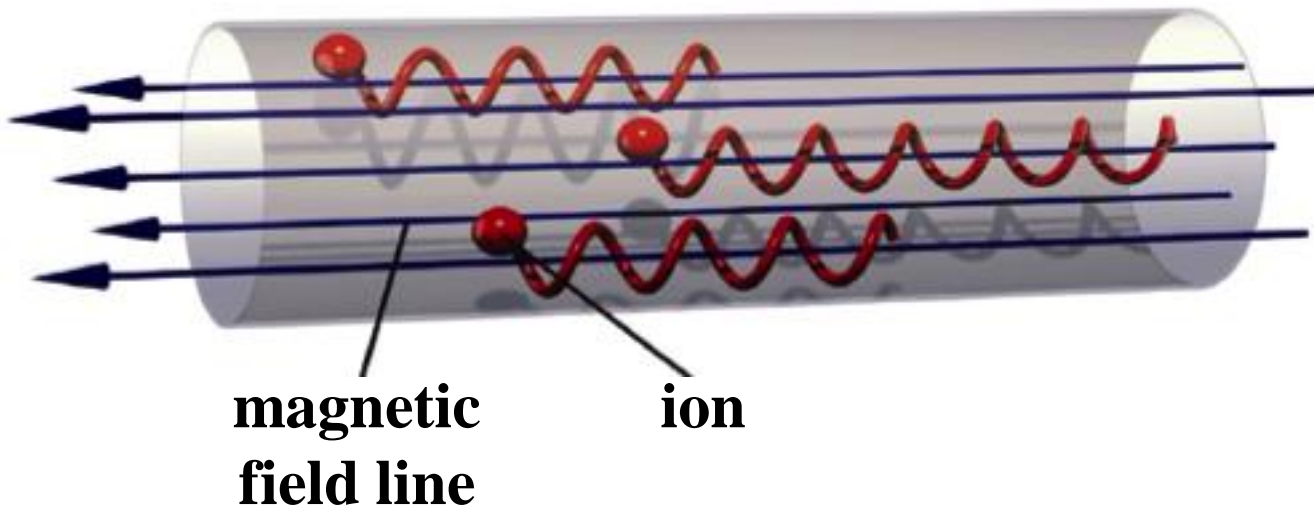
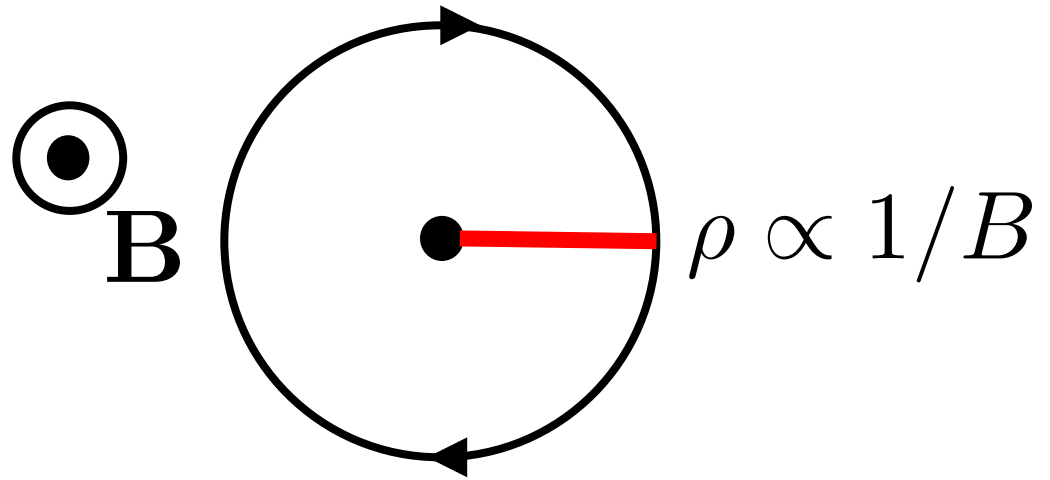
$$\text{Solve for required confinement time} \quad \longrightarrow \quad \tau_E \sim 1\text{s}$$

How to heat
the plasma
hotter than
the sun



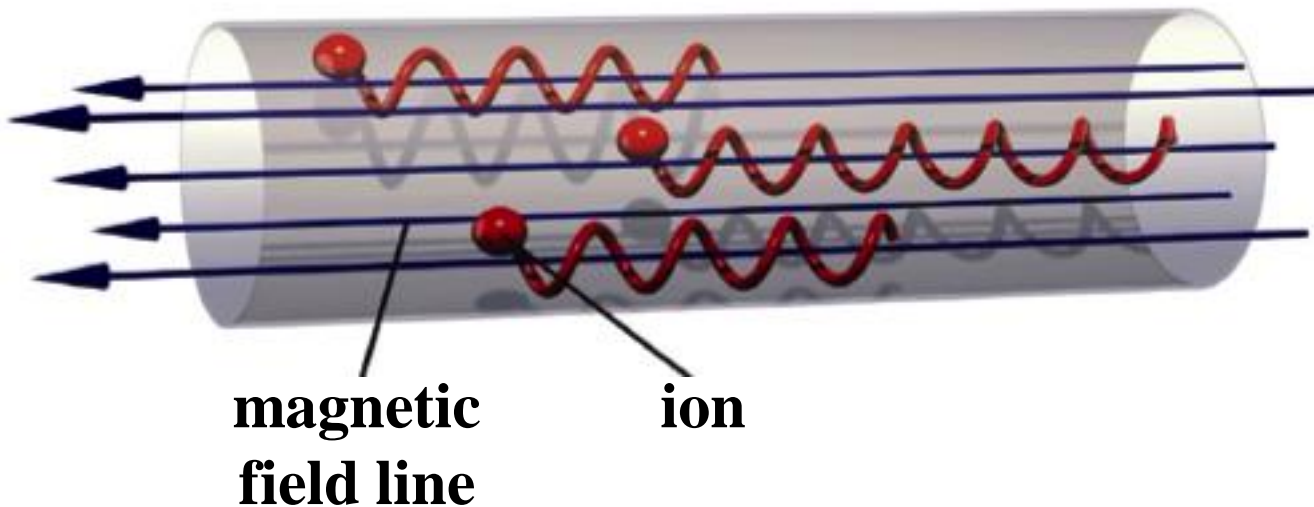
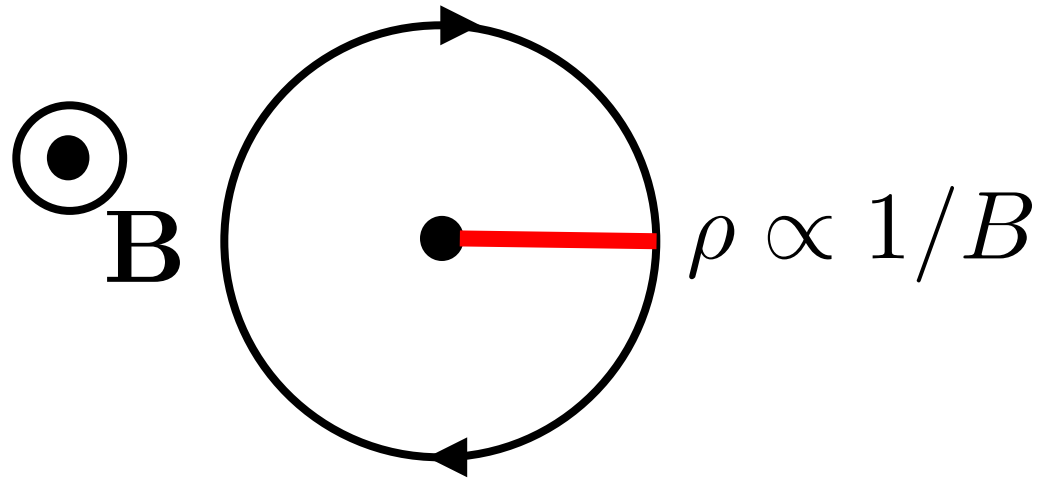
Basic concept of magnetic confinement

$$\mathbf{F} = \frac{e}{c} \mathbf{v} \times \mathbf{B}$$



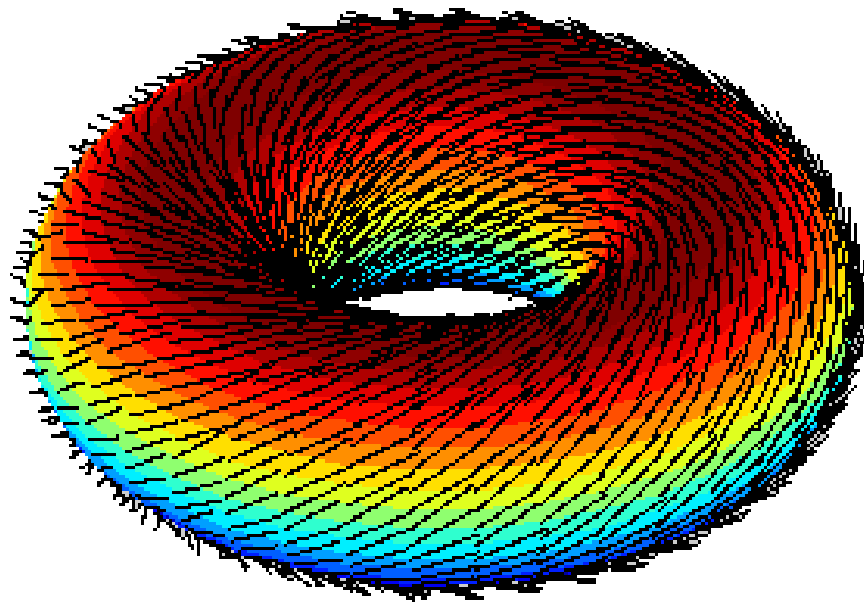
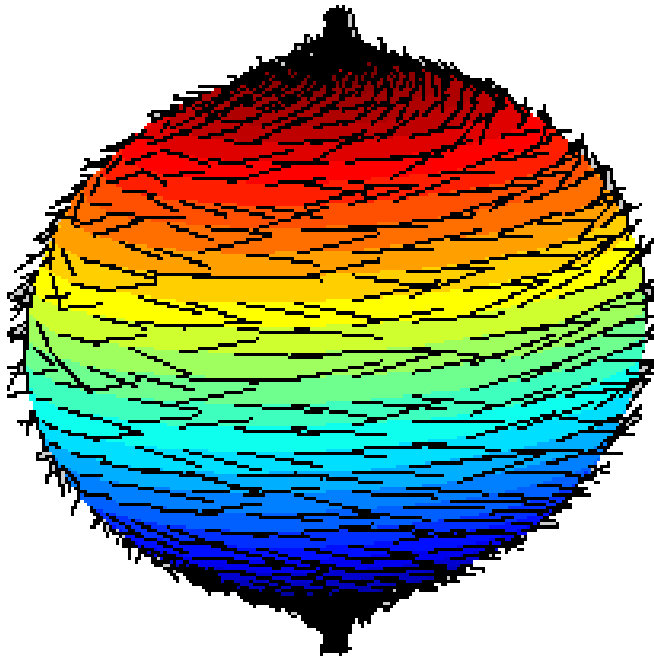
Stronger field = better confinement

$$\mathbf{F} = \frac{e}{c} \mathbf{v} \times \mathbf{B}$$



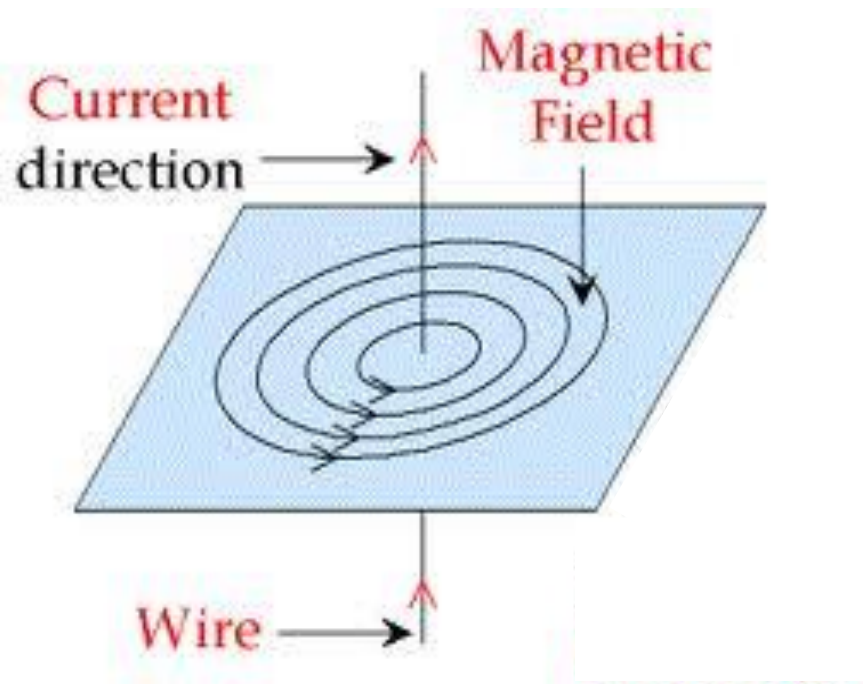
Why magnetic fusion devices look like doughnuts

- ‘Hairy ball theorem’ \rightarrow confined trajectories of vector field possible only for torii

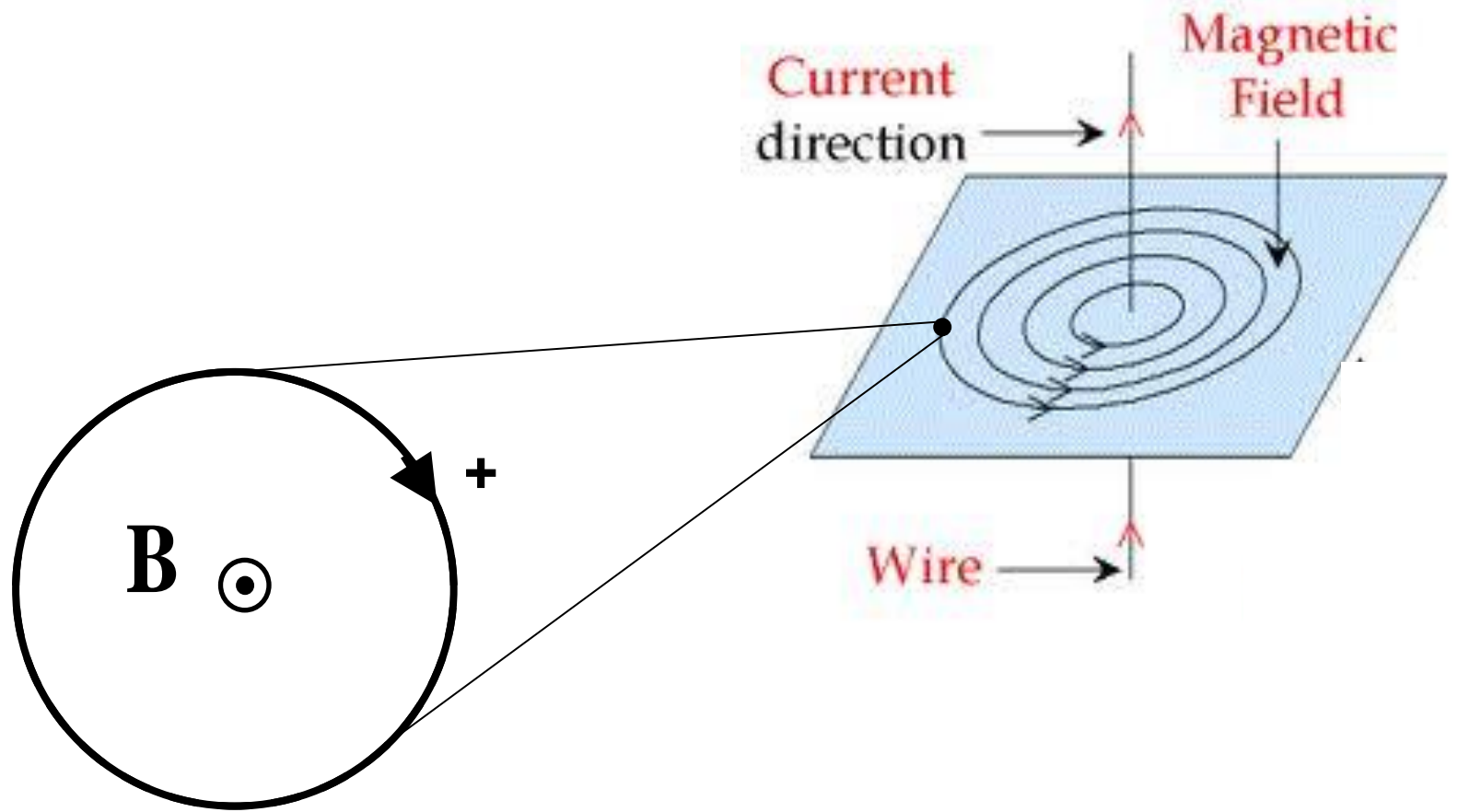


Confined field line trajectory

- Simplest idea is circles:



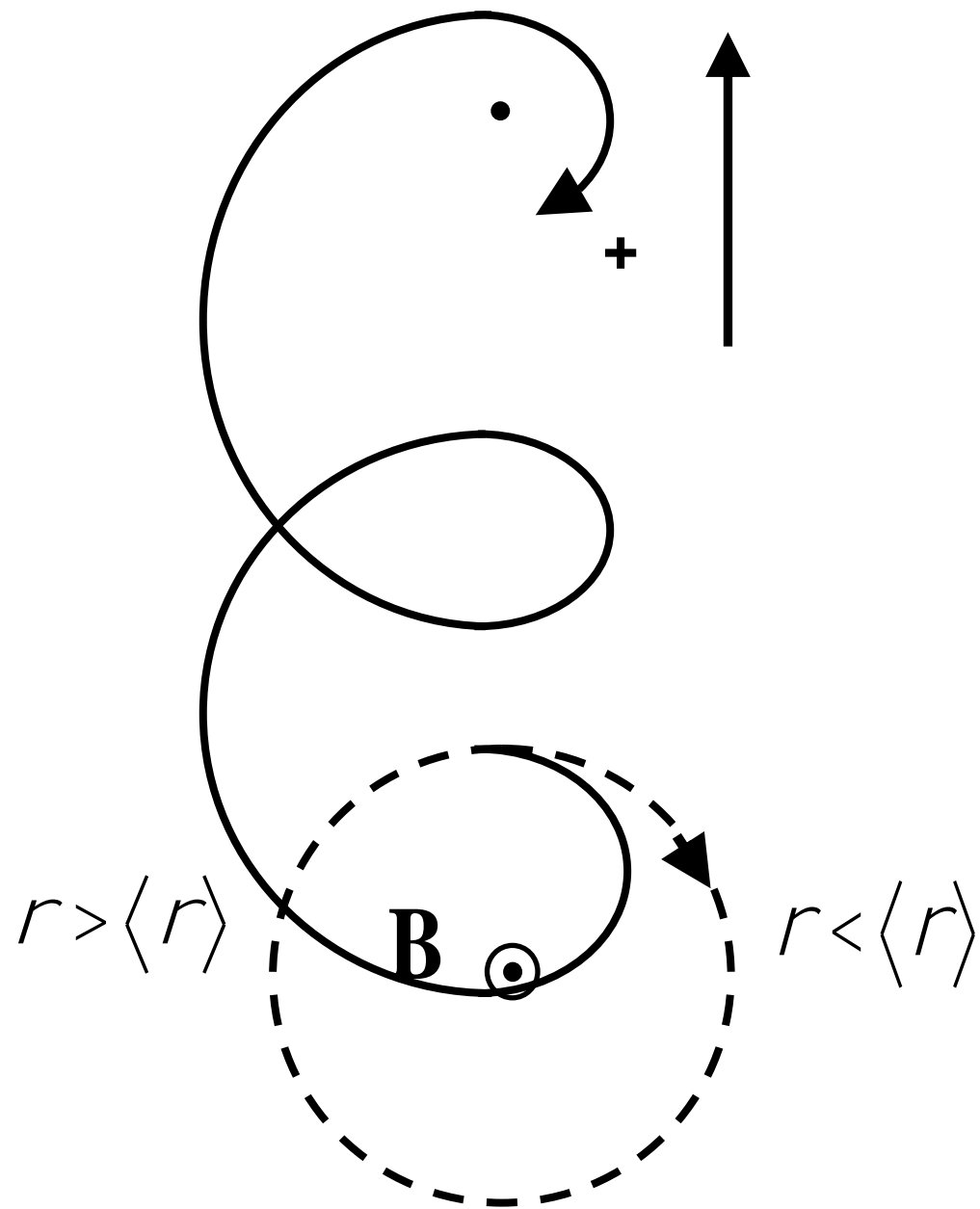
$$B = \frac{\mu_0 I}{2\rho R}$$



weaker field

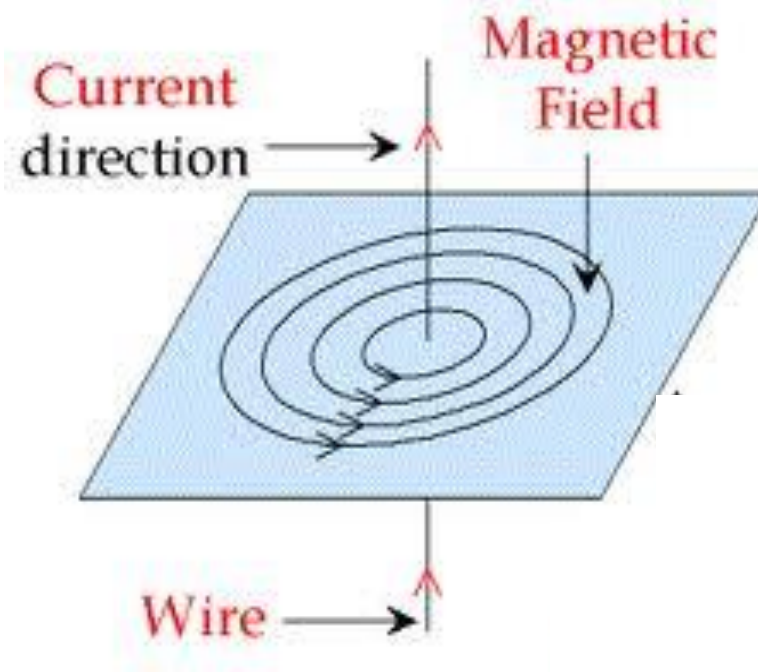
stronger field

$$\rho \propto 1/B$$

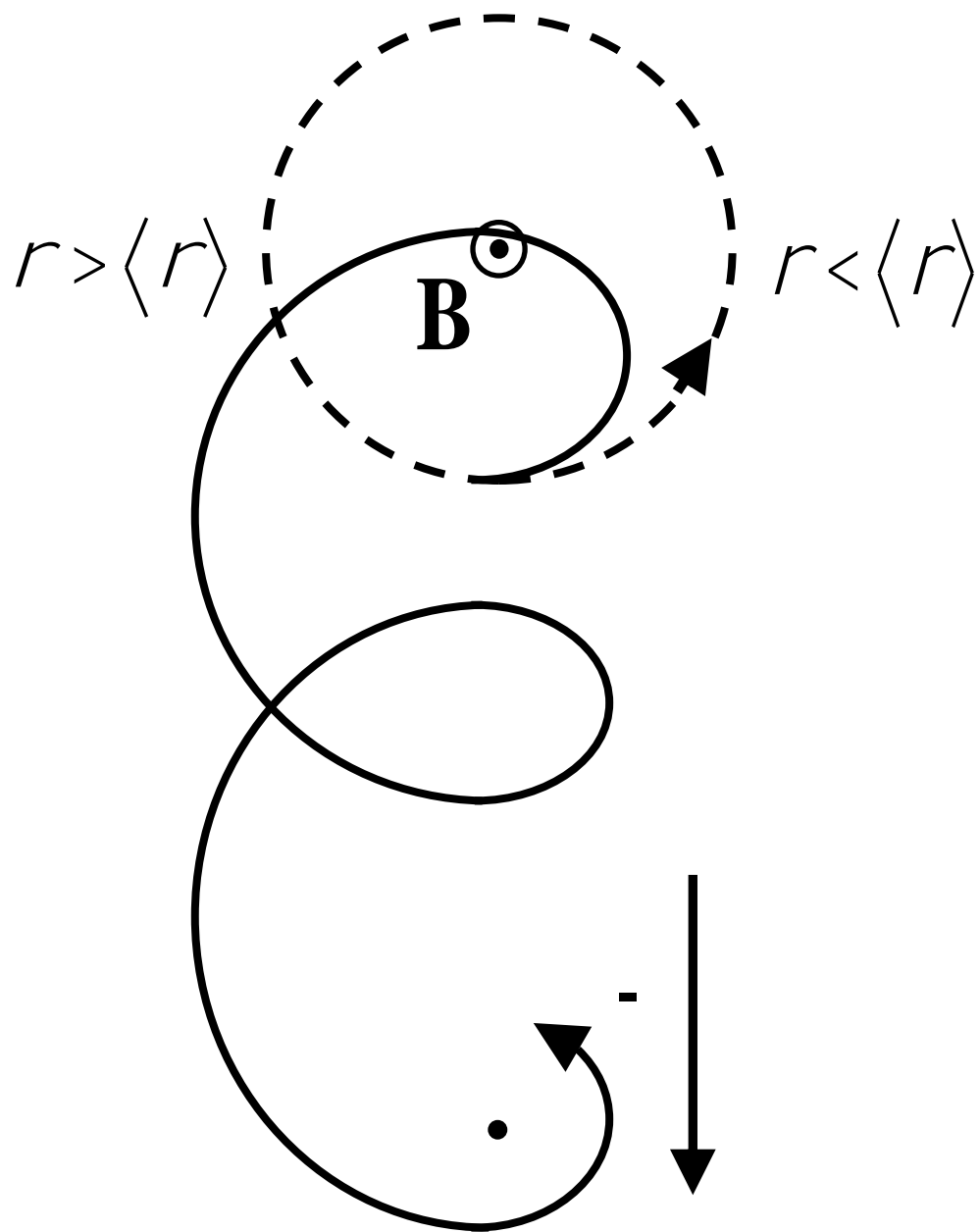


weaker field

stronger field

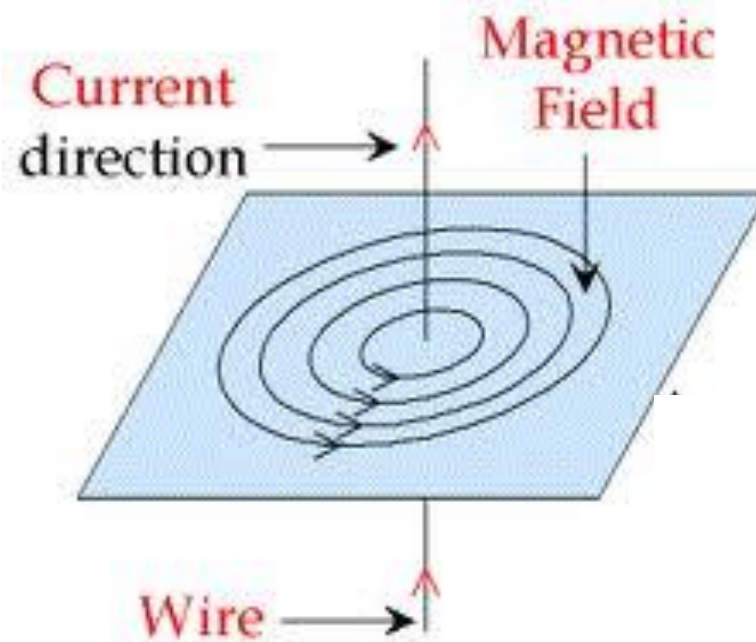


$$\rho \propto 1/B$$



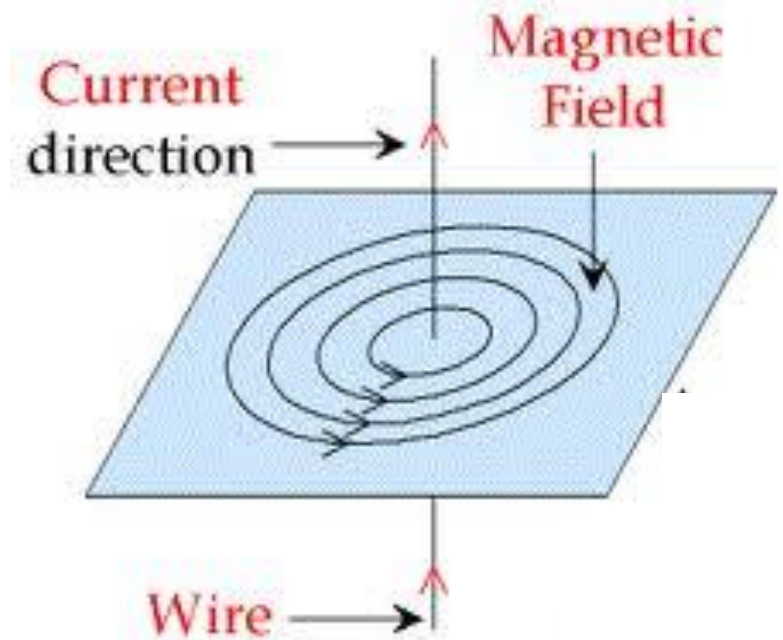
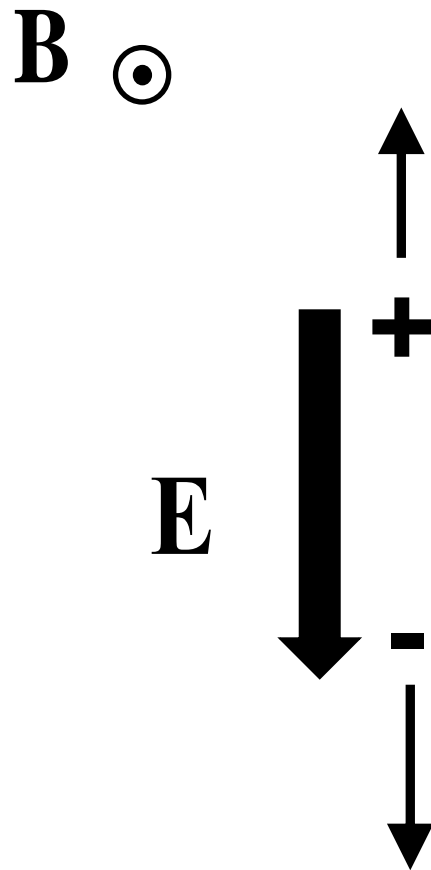
weaker field

stronger field

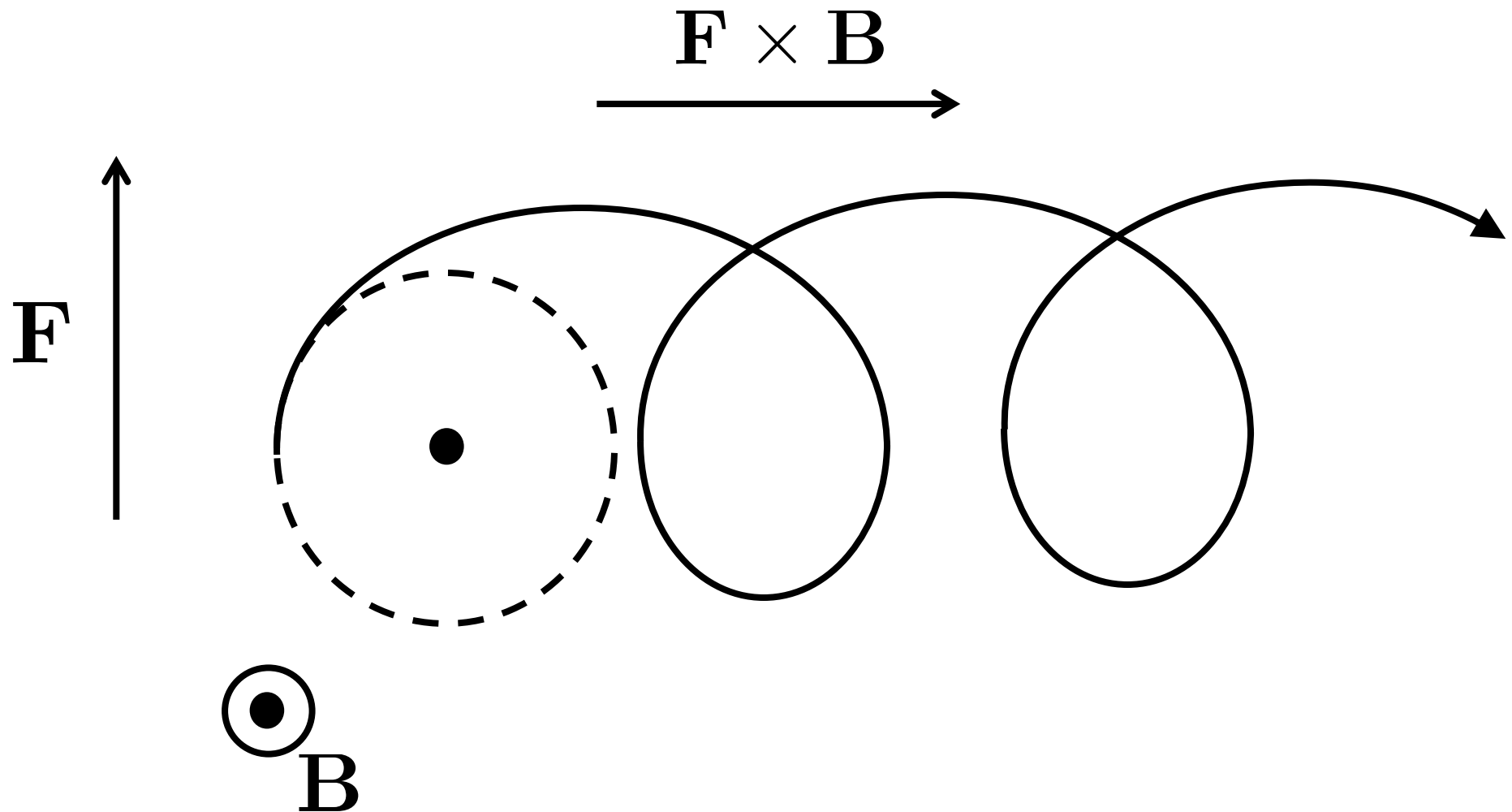


$$\rho \propto 1/B$$

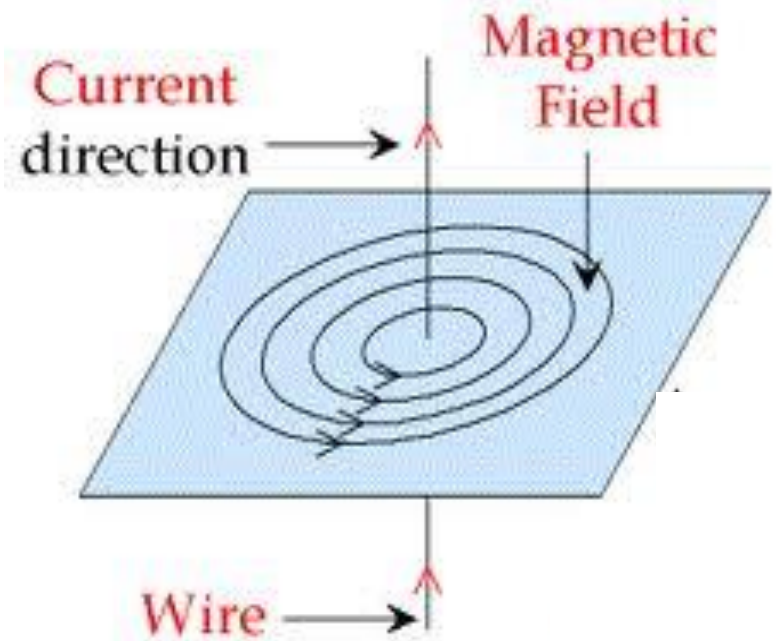
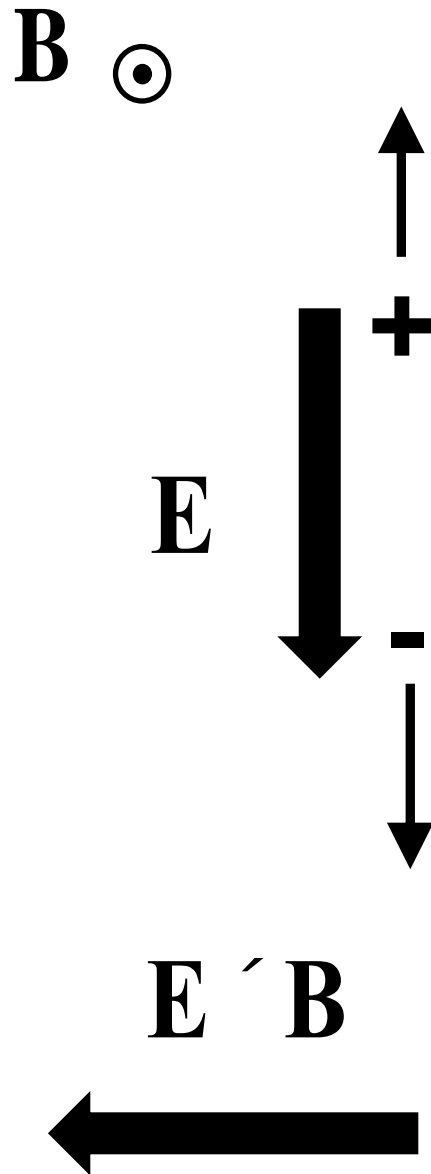
Charge separation gives electric field



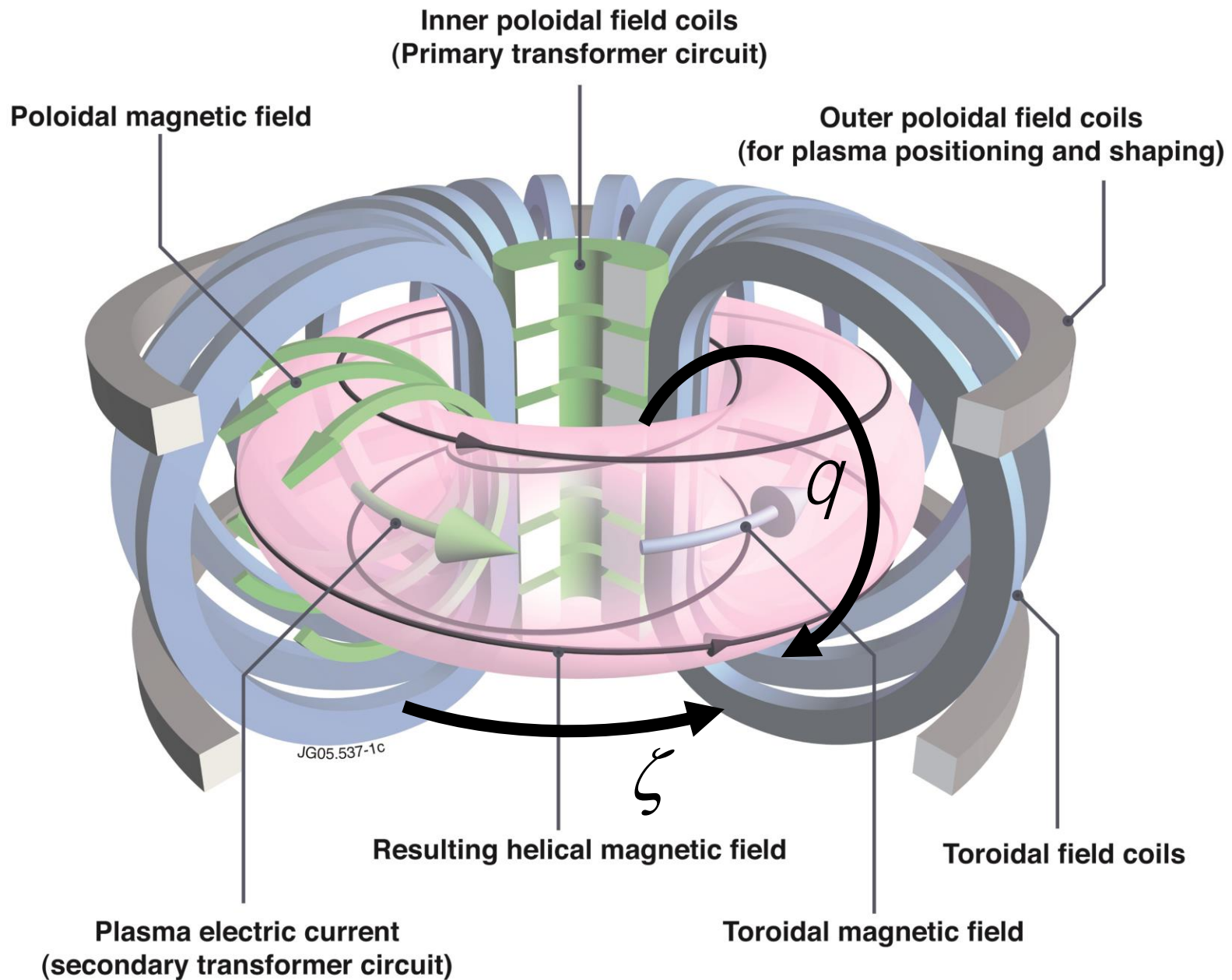
Magnetised plasma is different



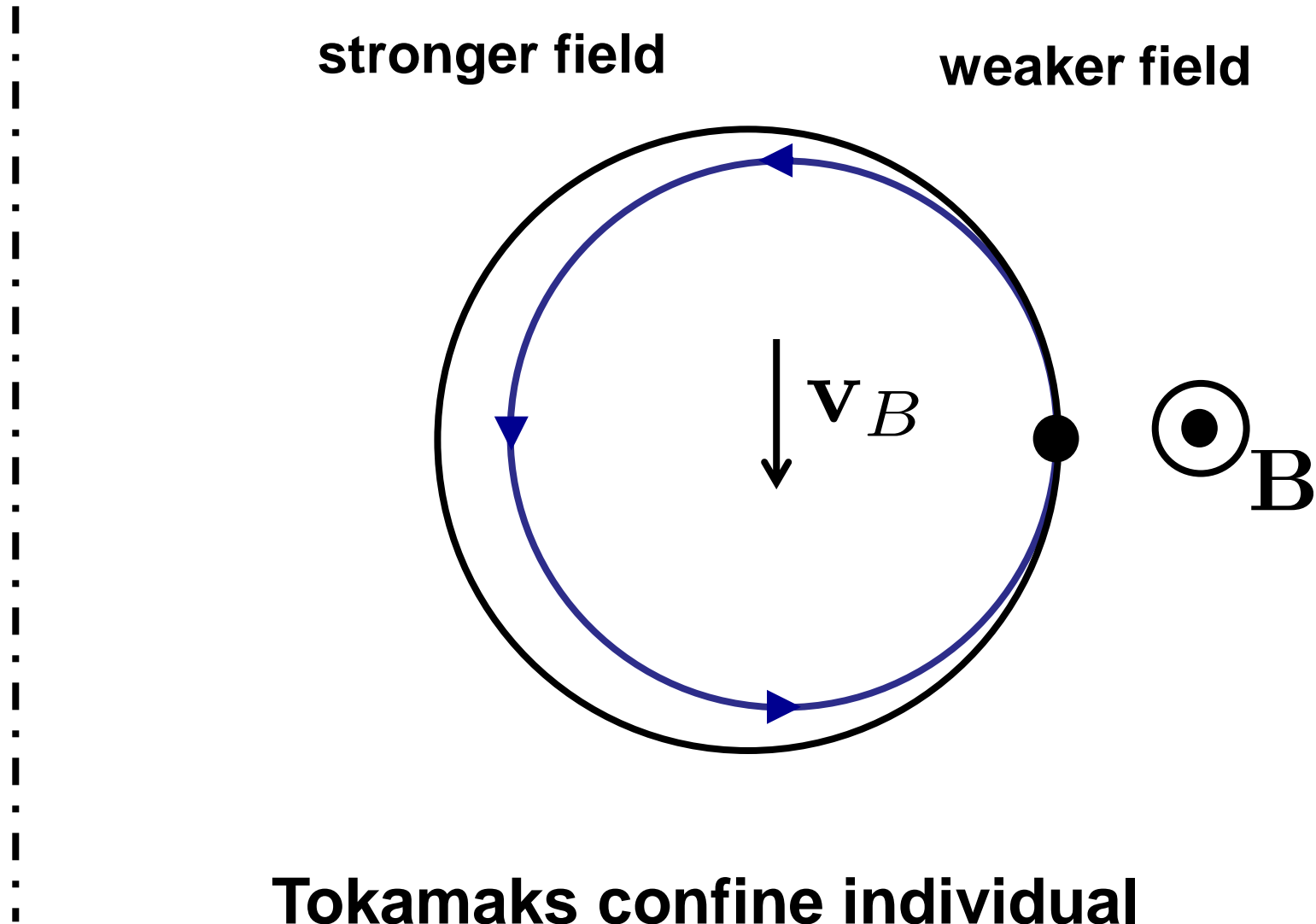
The simplest doughnut does not work



The solution? Add a twist

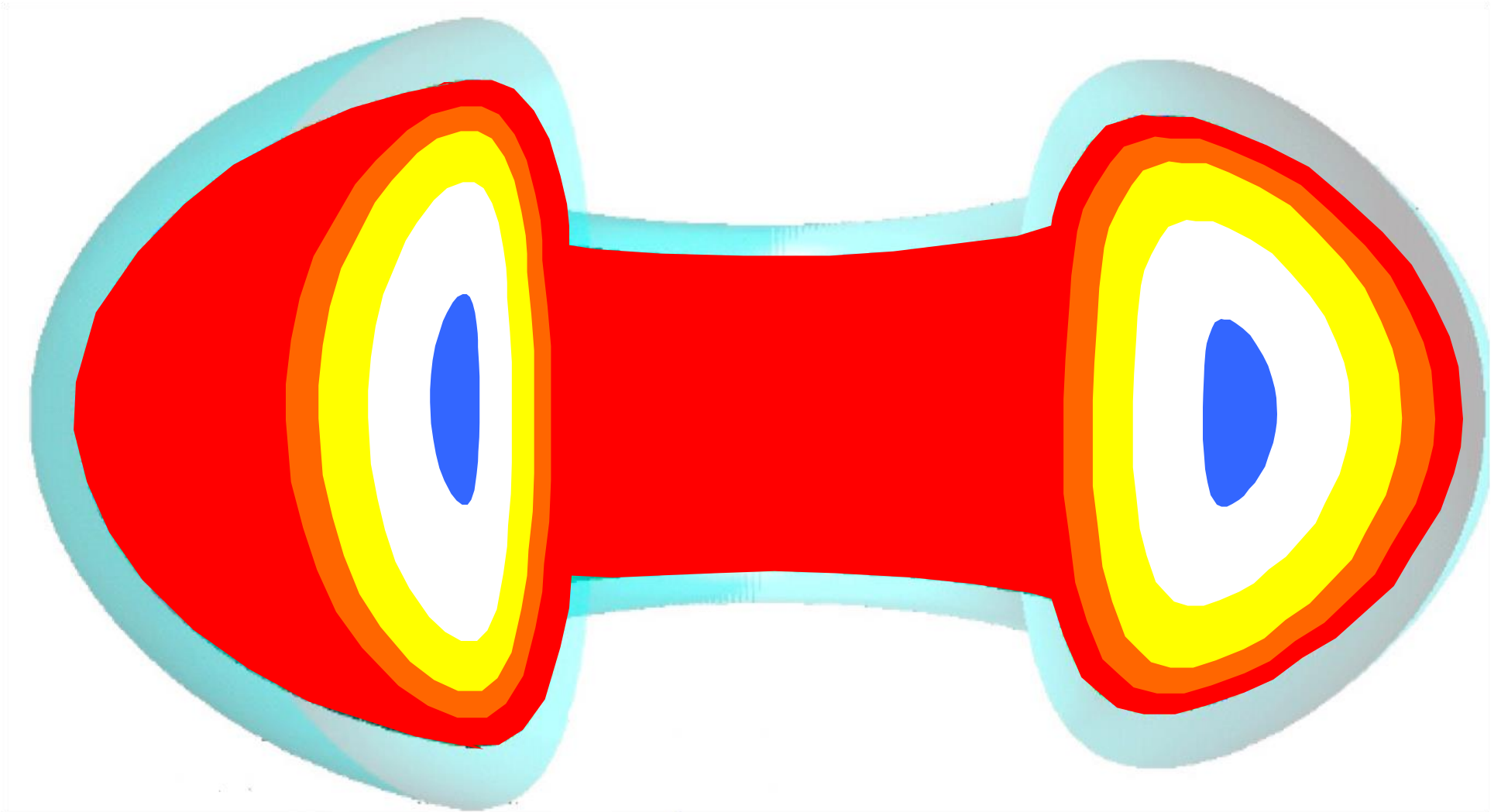


Magnetic drifts close, so no net drift



Tokamaks confine individual charged particles!

Magnetic confinement on toroidal surfaces



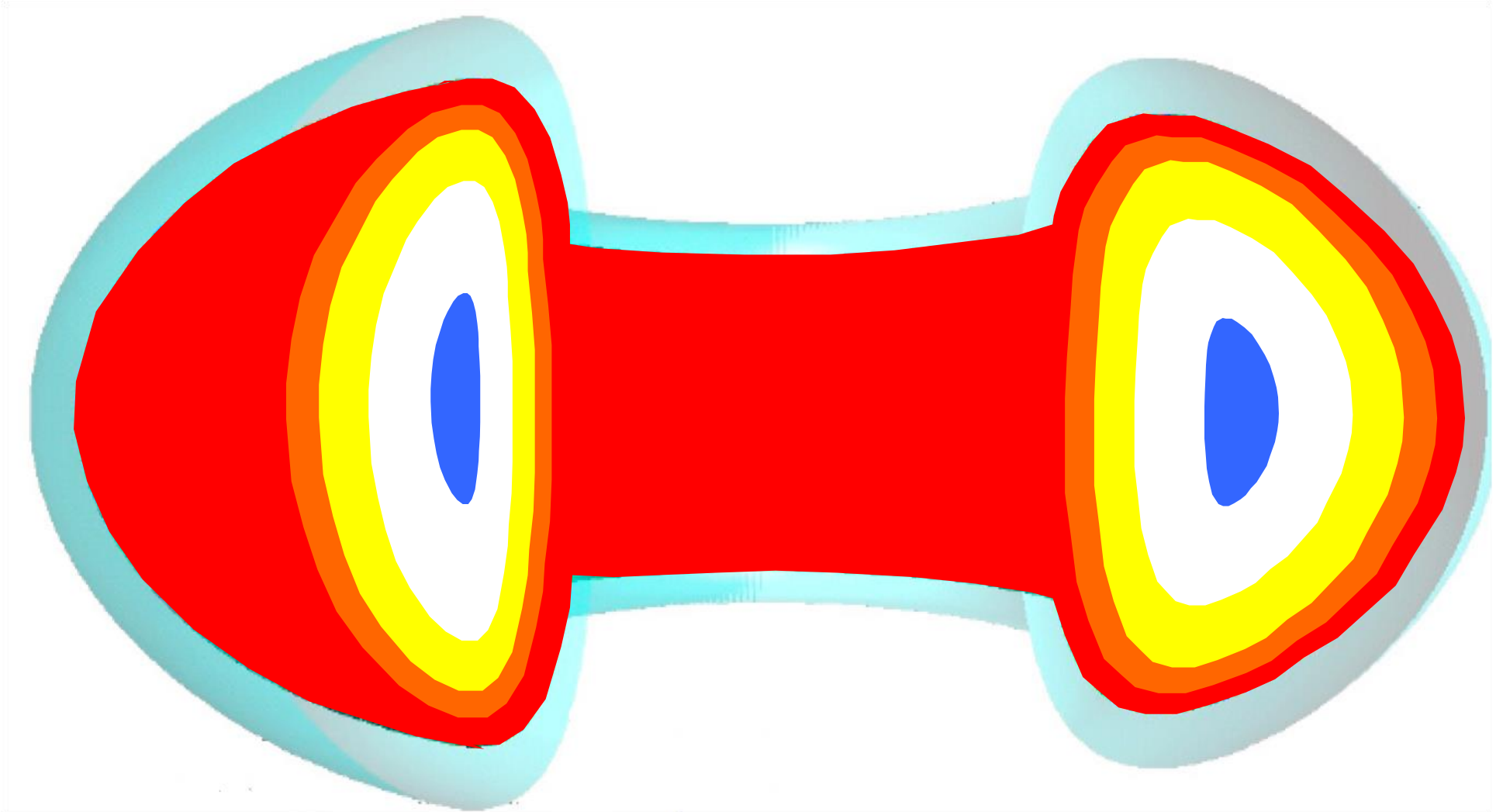
Nested contours of constant pressure/density/temperature

Challenges and opportunities ahead

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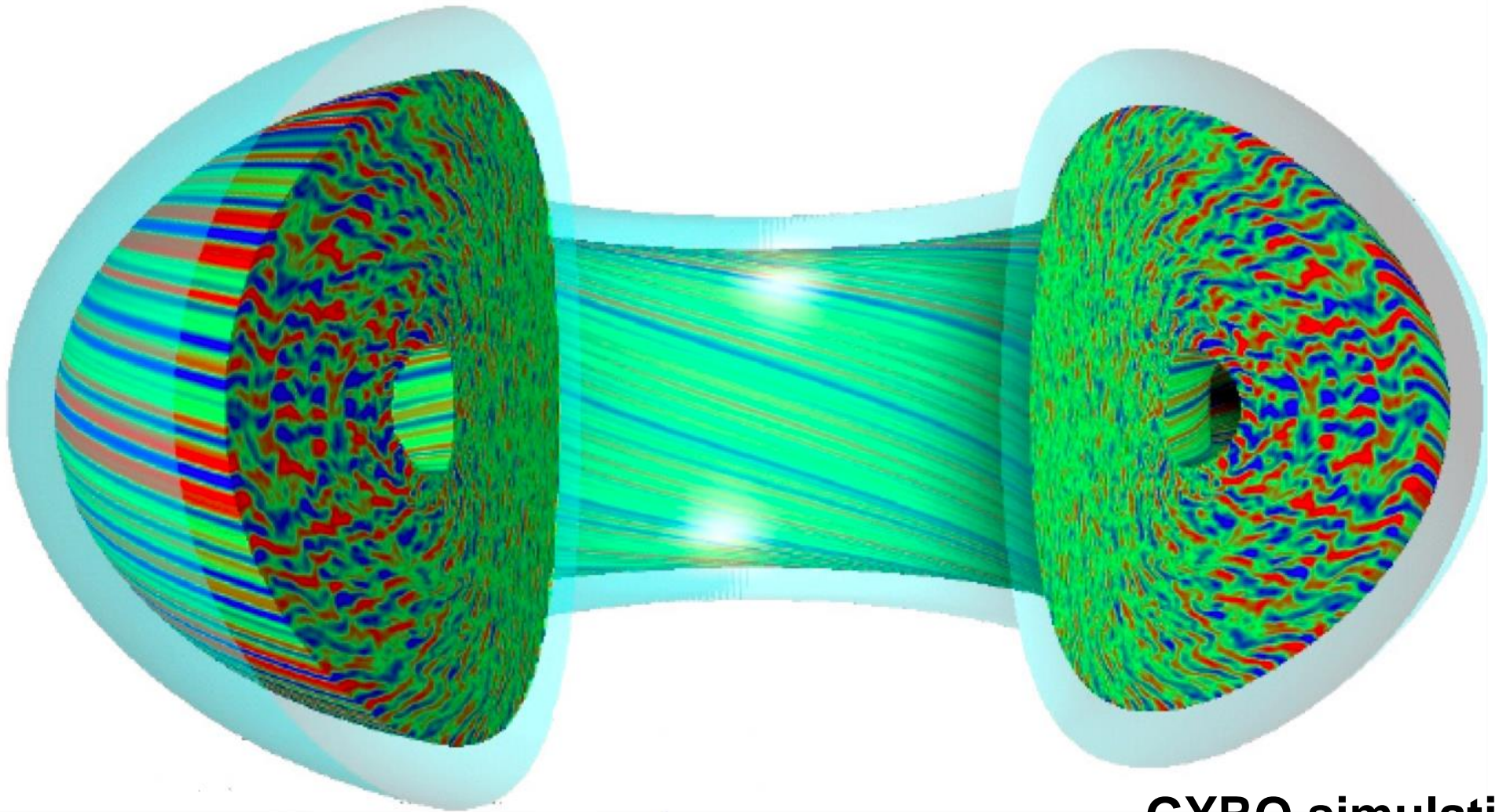
Fast camera image of MAST plasma

Magnetic confinement on toroidal surfaces



Nested contours of constant pressure/density/temperature

Magnetic confinement on toroidal surfaces...almost



GYRO simulation

Turbulent density fluctuations

Words of encouragement

Sir Horace Lamb
(1904)

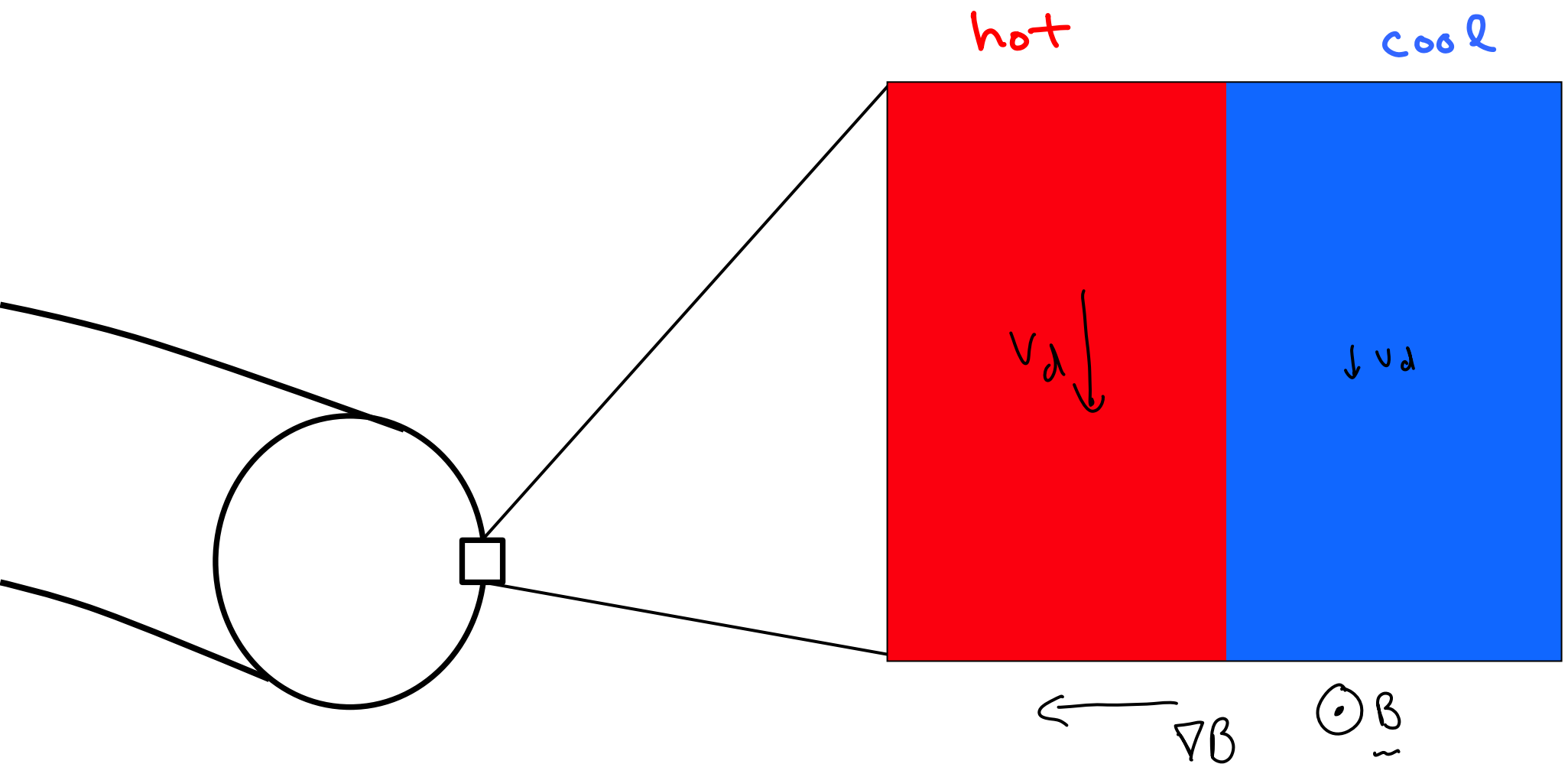


“When I meet God, I am going to ask him two questions: Why relativity? And why turbulence? I really believe he will have an answer for the first.”

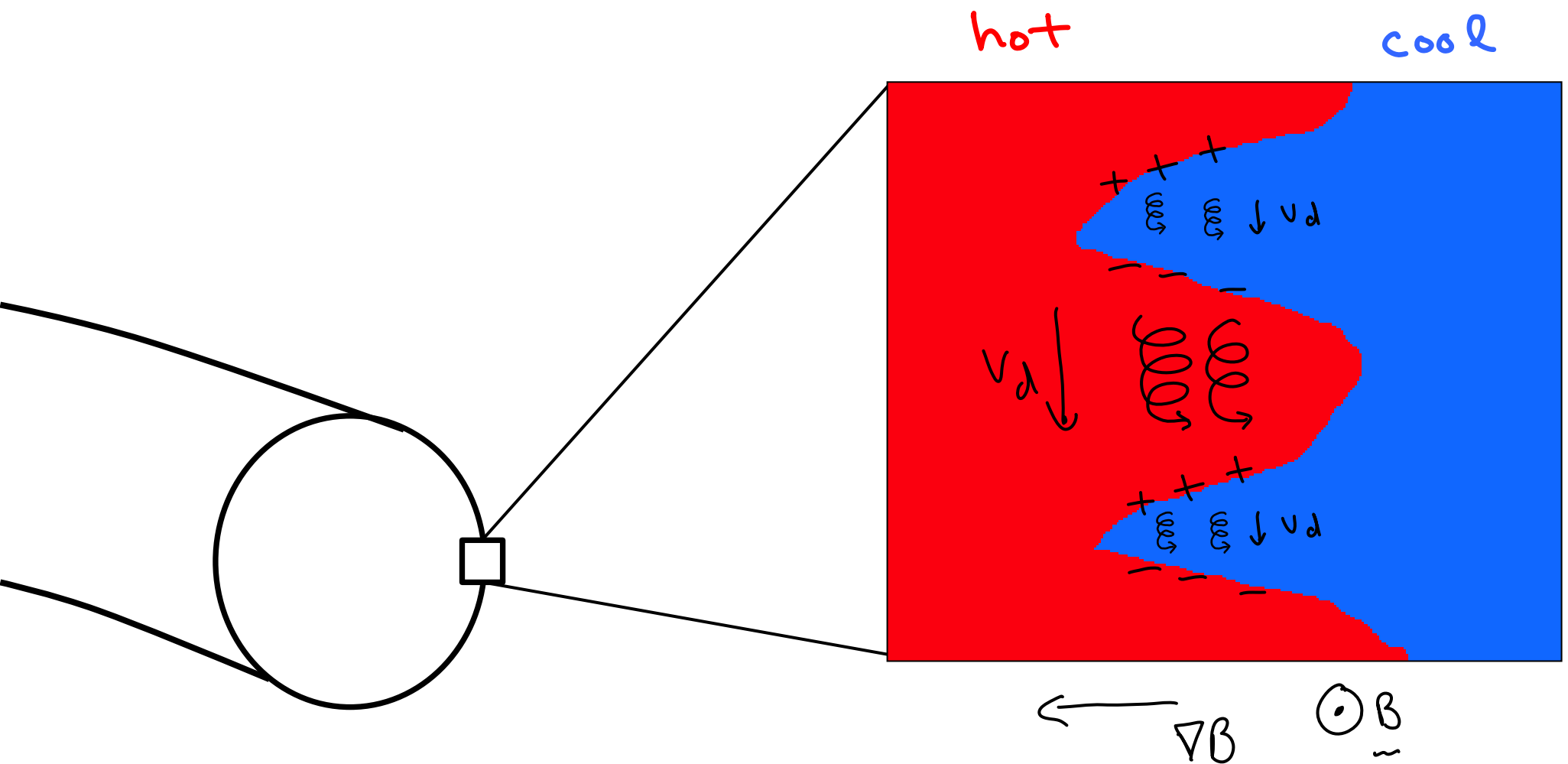
Werner Heisenberg (1933),
German Federal Archives



What drives the turbulence?

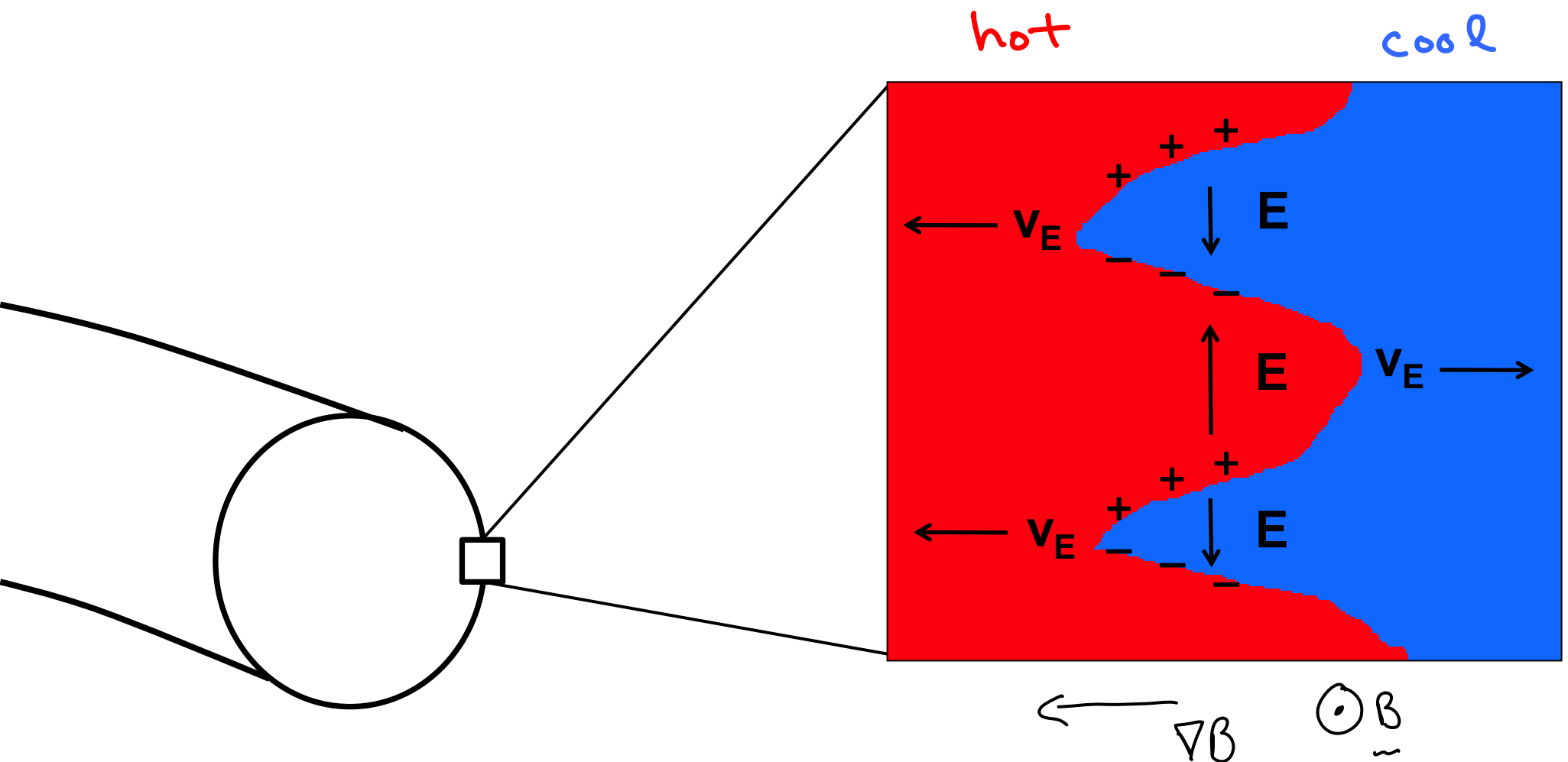


What drives the turbulence?



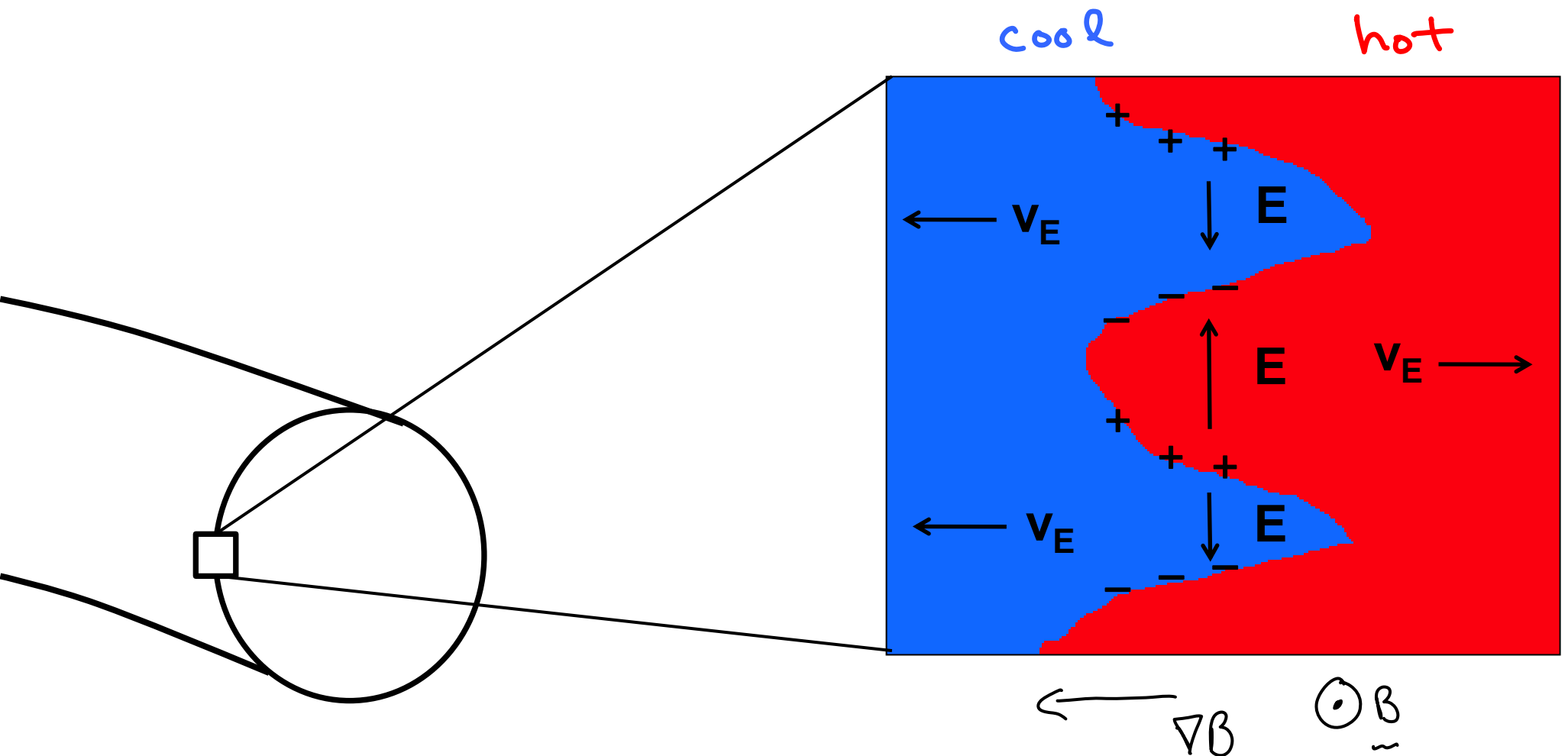
What drives the turbulence?

$$\mathbf{v}_E = \frac{c}{B^2} \mathbf{E} \times \mathbf{B}$$

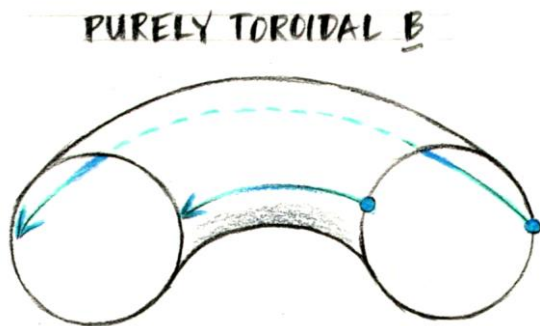


No drive on inside

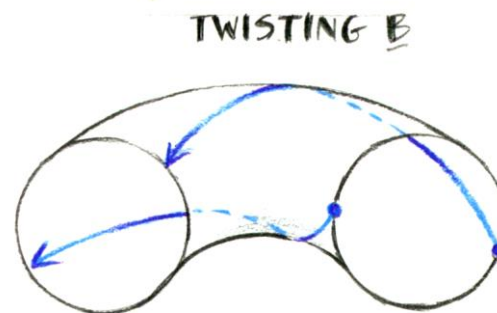
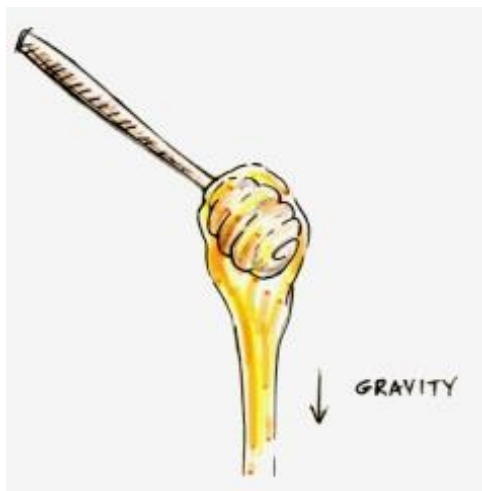
$$\mathbf{v}_E = \frac{c}{B^2} \mathbf{E} \times \mathbf{B}$$



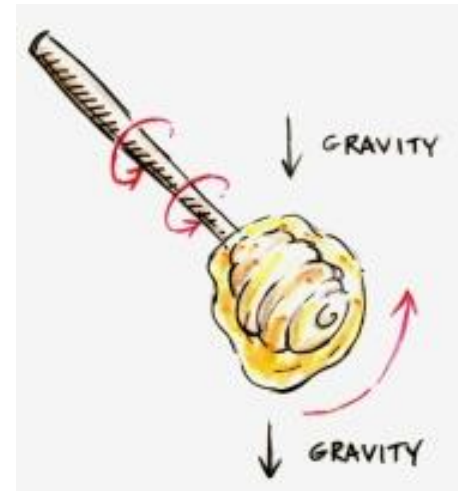
Competition gives critical gradient



Unstable



Stable

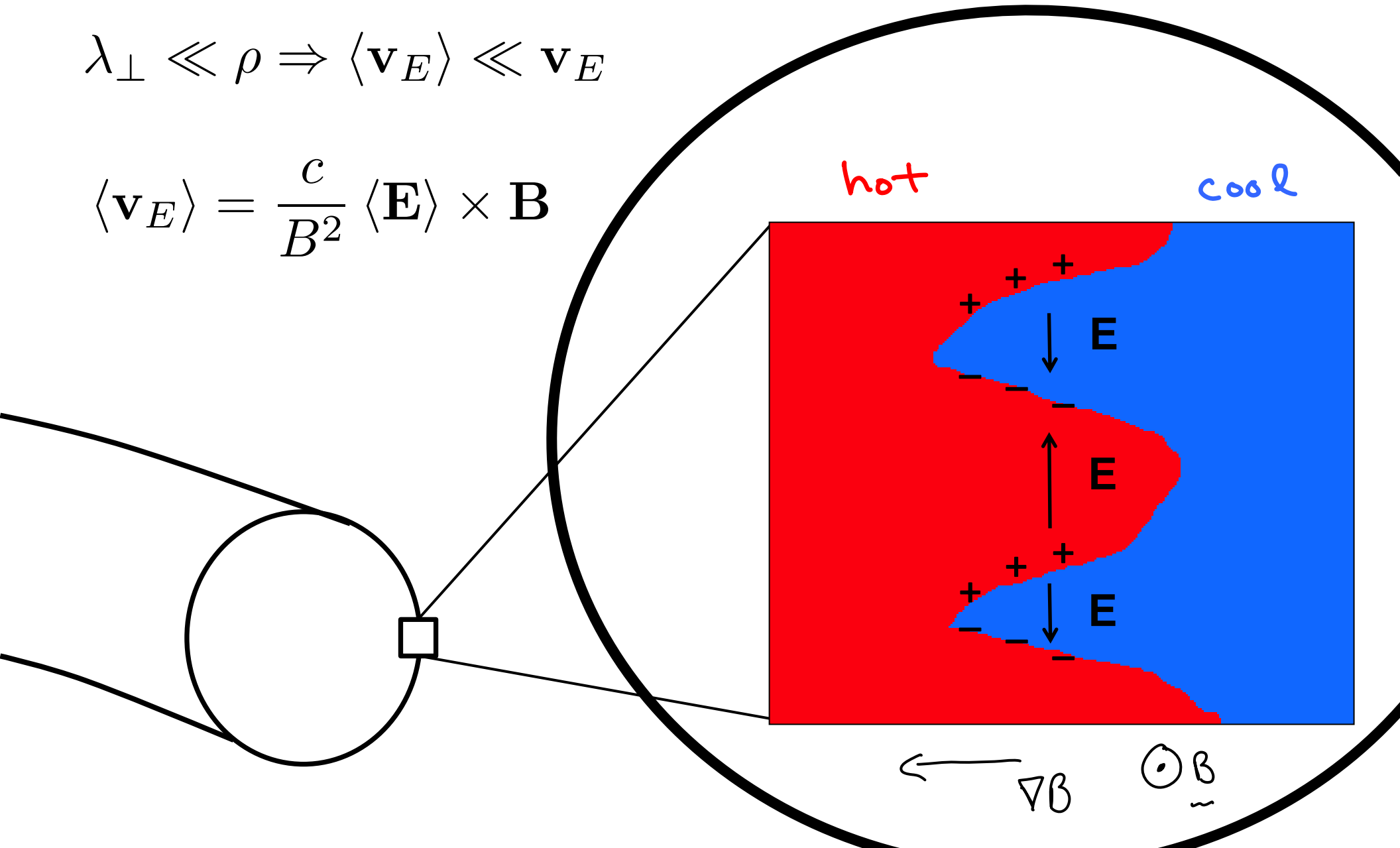


No turbulence below some critical temperature gradient

Eddy size is gyro-radius

$$\lambda_{\perp} \ll \rho \Rightarrow \langle \mathbf{v}_E \rangle \ll \mathbf{v}_E$$

$$\langle \mathbf{v}_E \rangle = \frac{c}{B^2} \langle \mathbf{E} \rangle \times \mathbf{B}$$

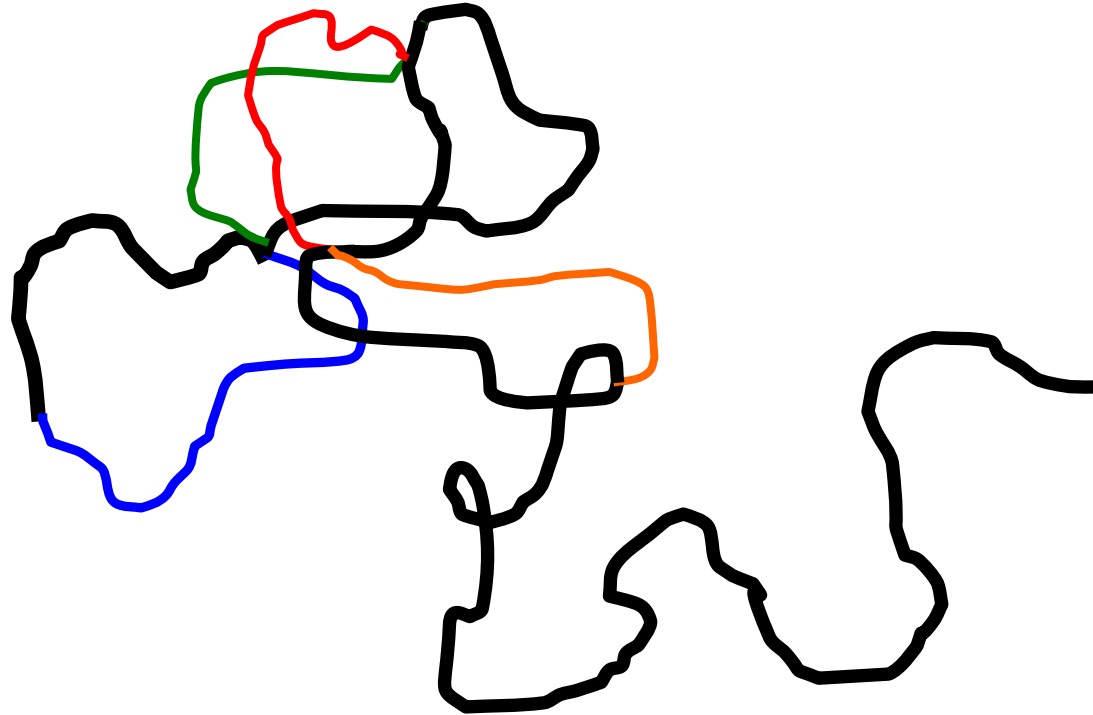


DIII-D Shot 121717

GYRO Simulation

Cray X1E, 256 MSPs

Turbulence and random walks

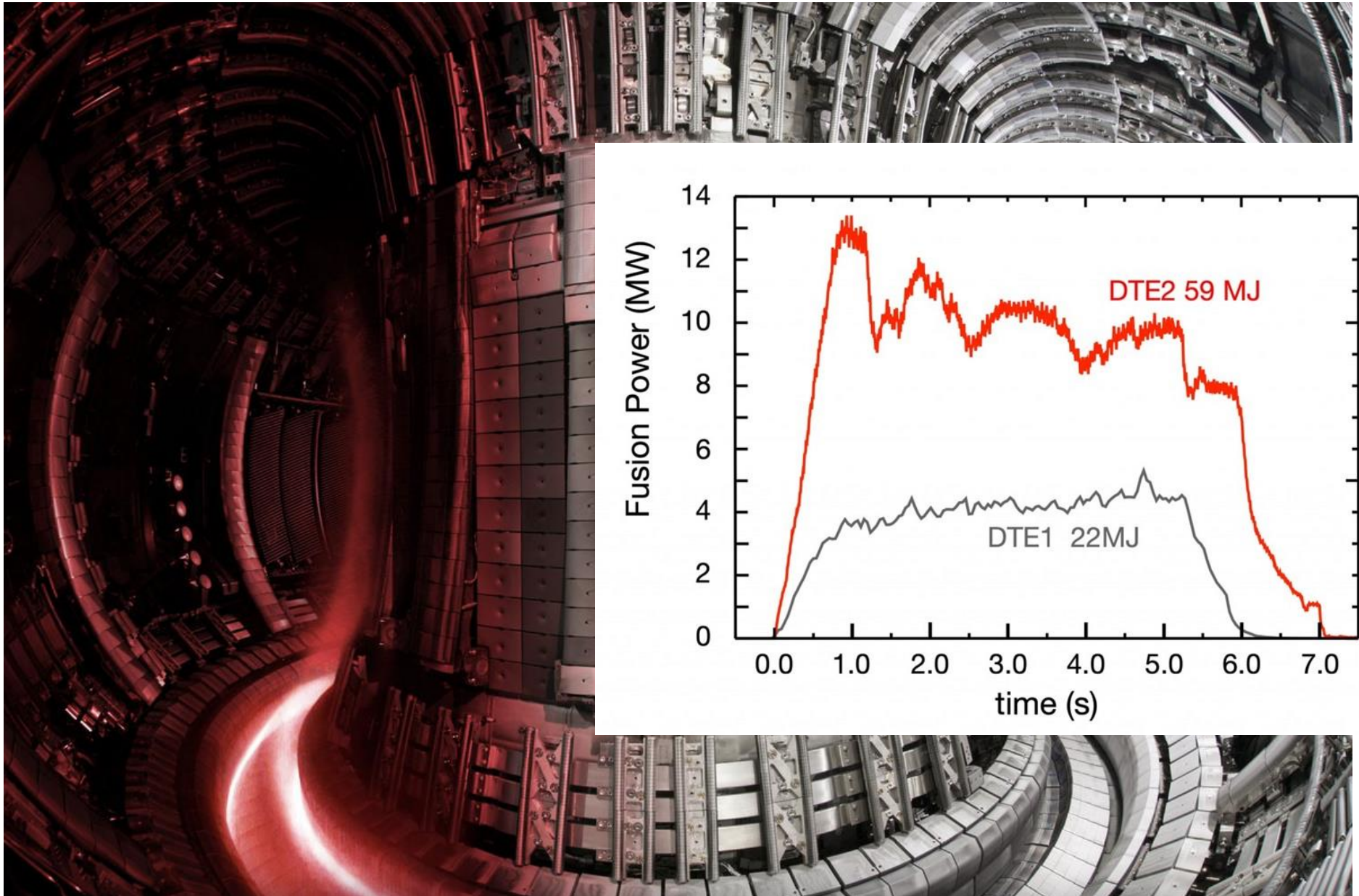


**Random walk: (time to move distance L) =
(time per step) x (L/d) steps**

L = system size, d = gyroradius, time per step = turbulence time

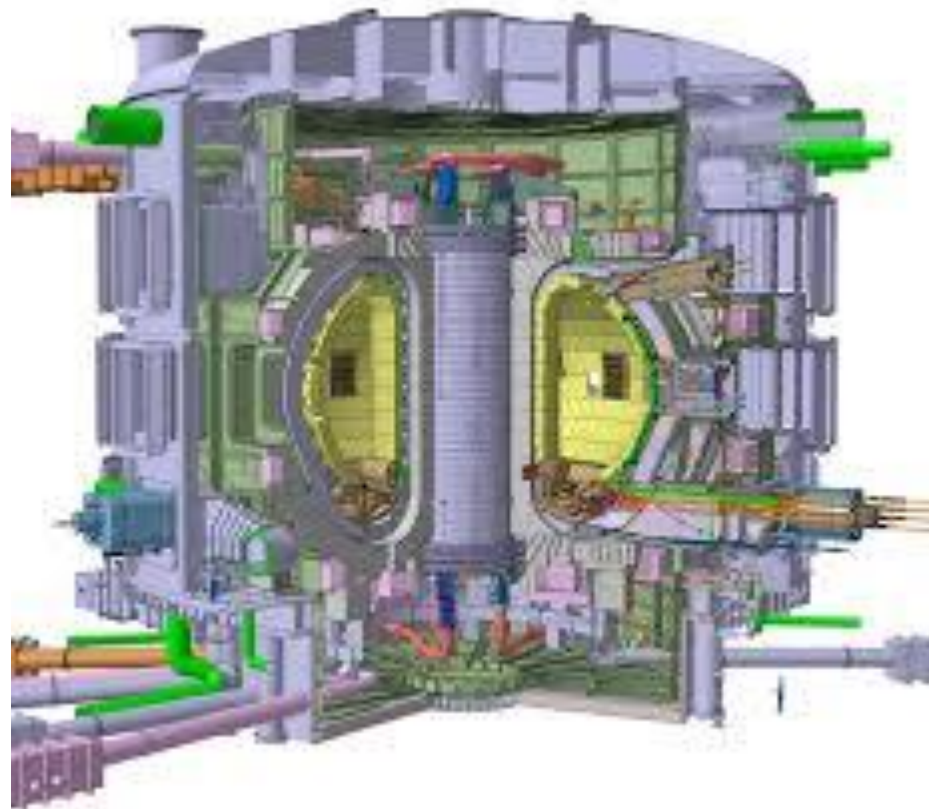
$$\longrightarrow \tau_E \sim 1 \text{ s}$$

We are close to the required confinement time, but not there yet



Different reactor approaches

- Make the plasma volume bigger
 - Pros: Achieve fusion temperatures without exceeding critical gradient that drives instability
 - Cons: bigger = more expensive; high heat loads on walls



ITER

ITER first plasma 2026

Largest scientific
experiment ever built

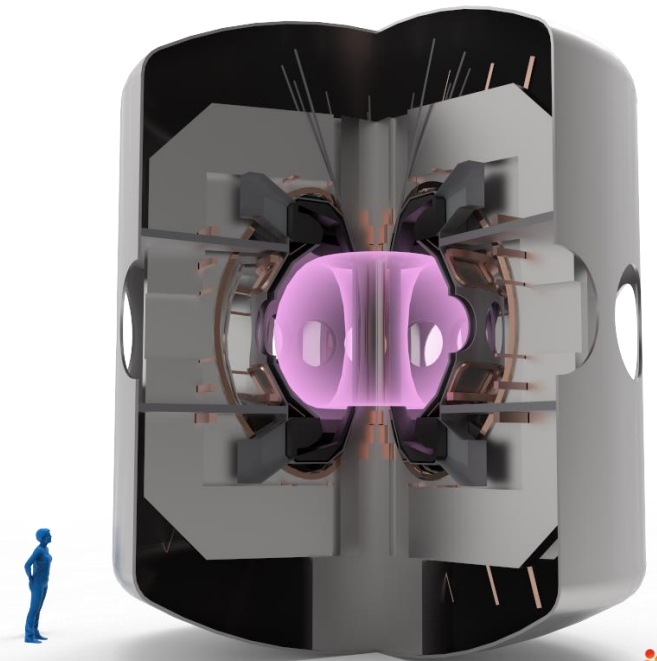
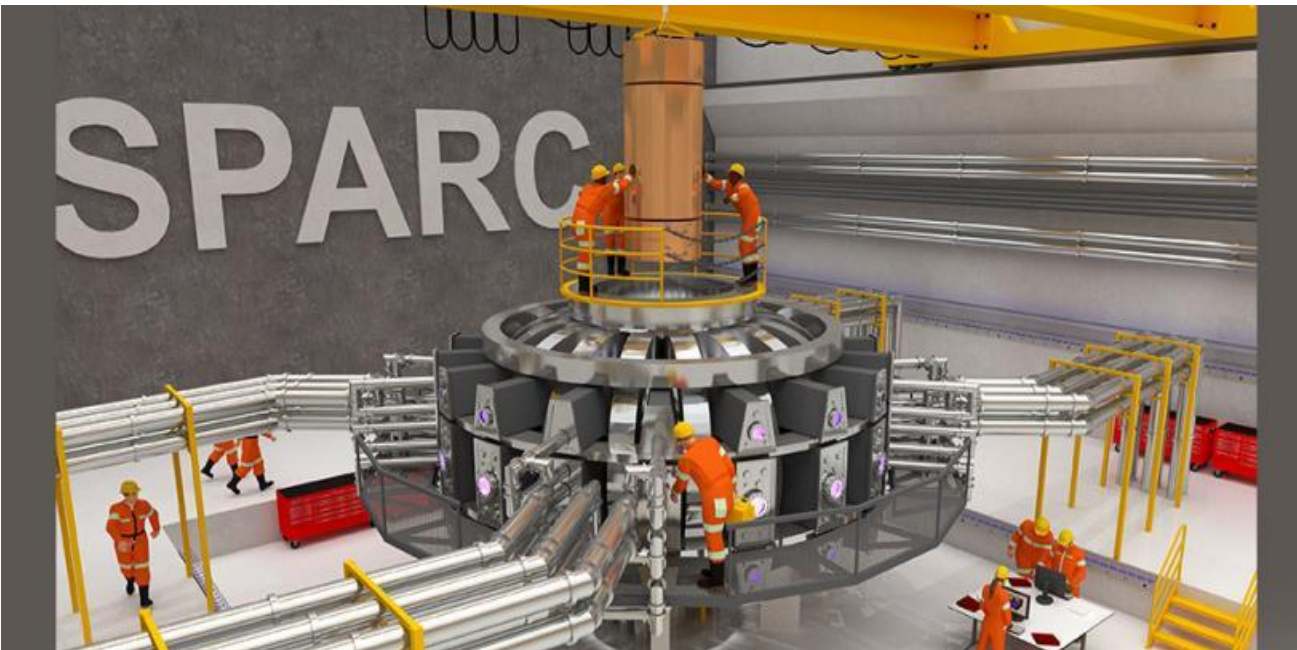
Construction towards 1st
plasma is 73% complete

May 2021



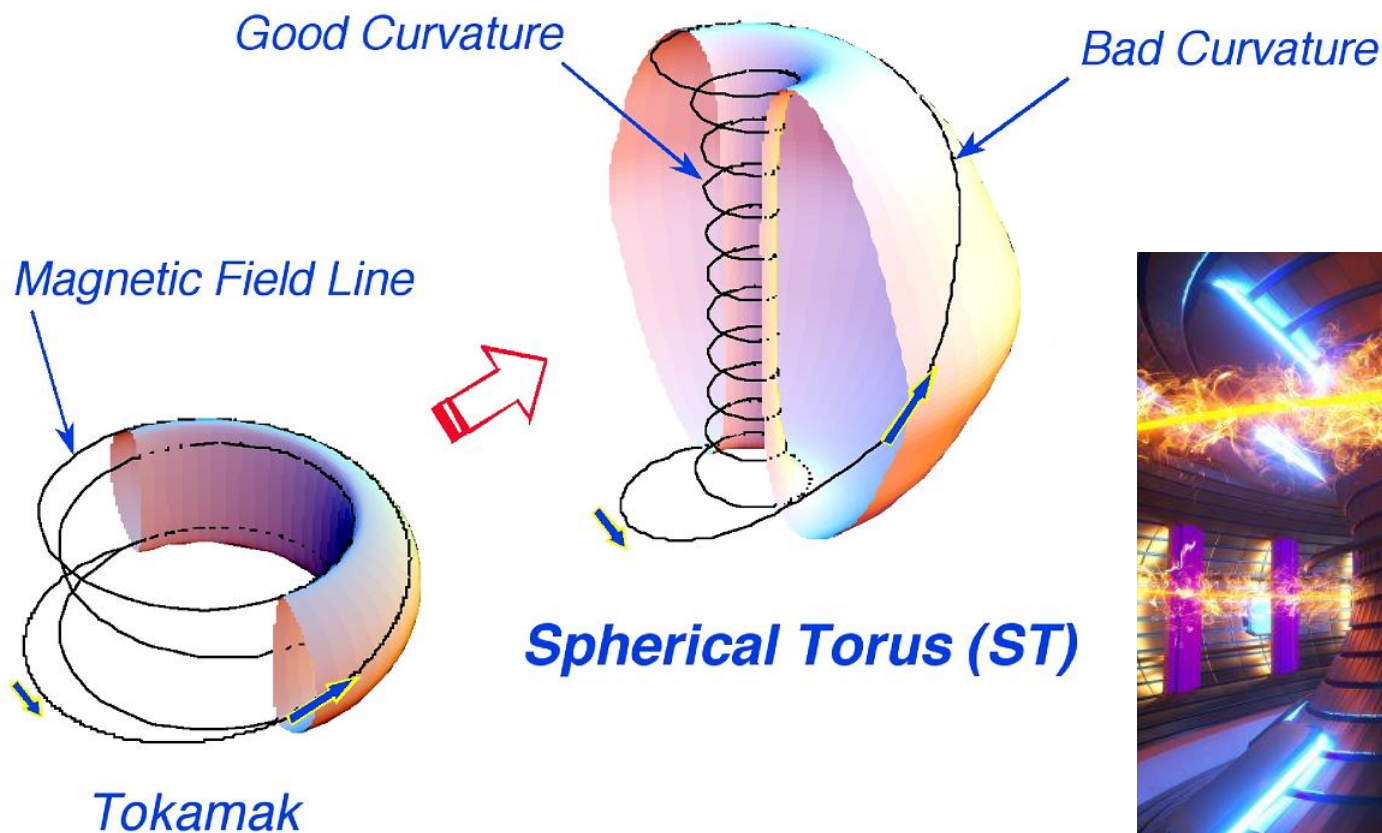
Different reactor approaches

- Make the magnetic field stronger
 - Pro: increased confinement time (smaller eddies = smaller device = cheaper)
 - Con: technology in development; high stresses

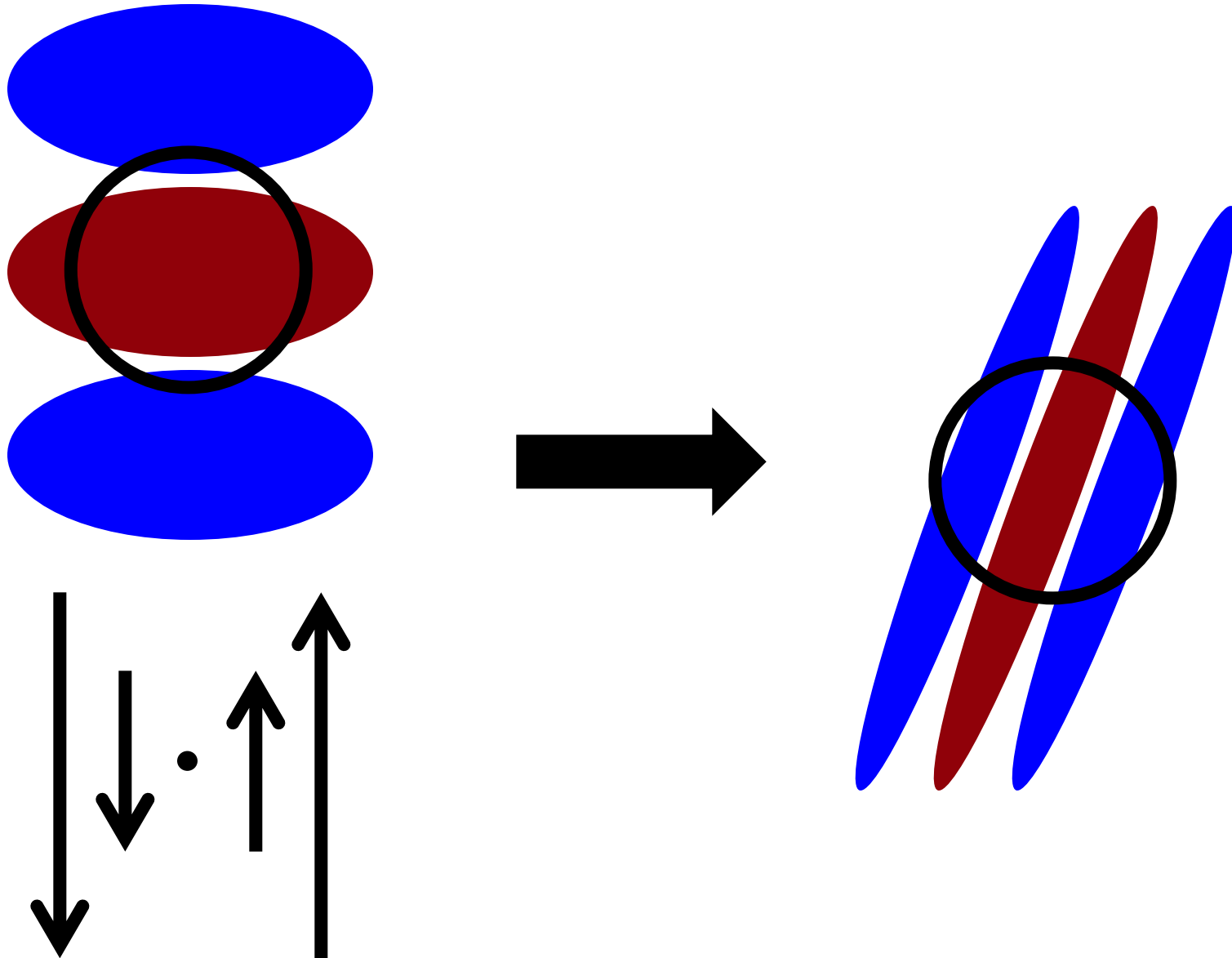


Different reactor approaches

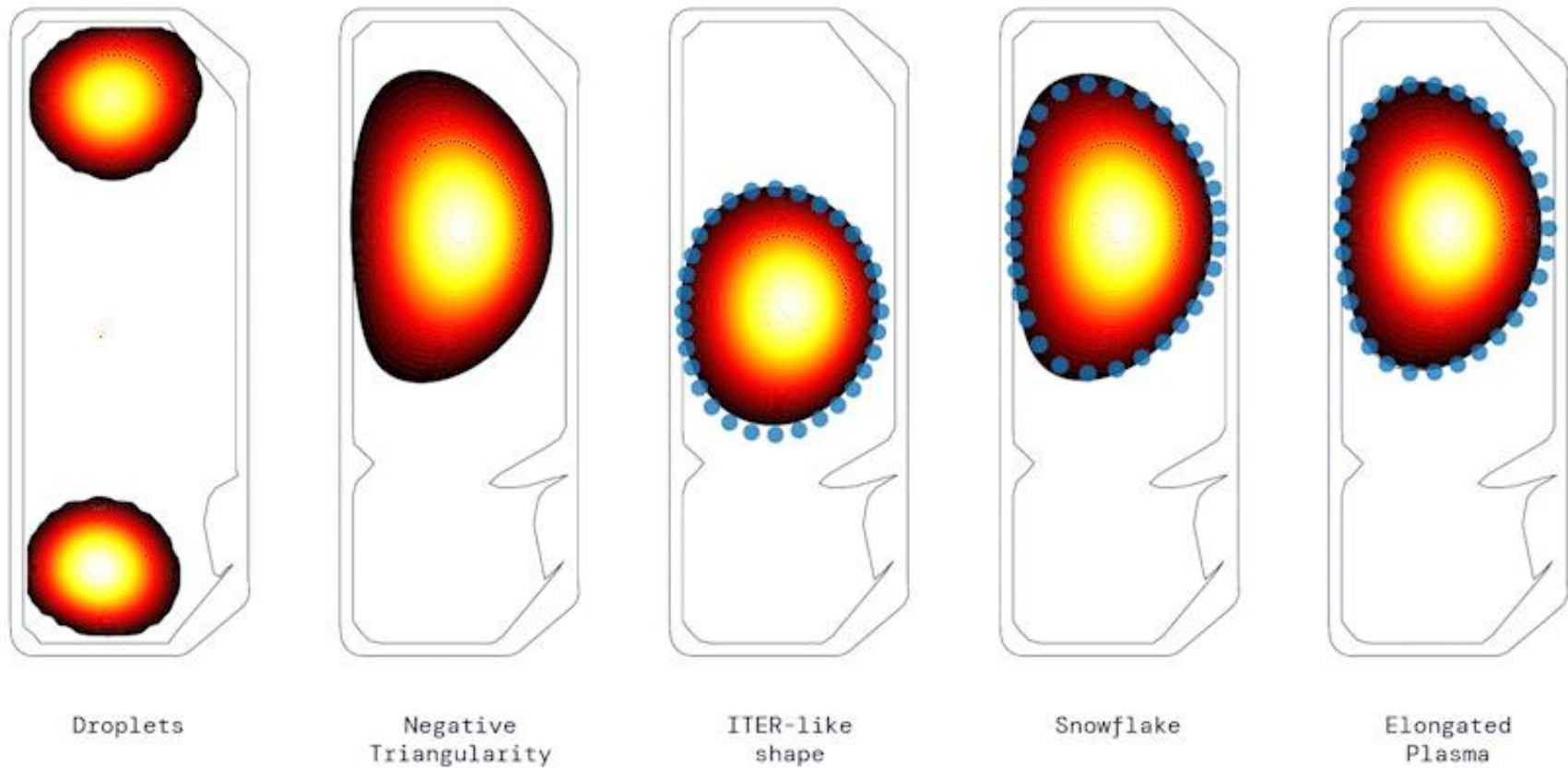
- Reduce the aspect ratio
 - Pro: cheaper, better stability/confinement
 - Con: engineering constraints (shielding + heat loads)



Turbulence suppression via, e.g., flow shear

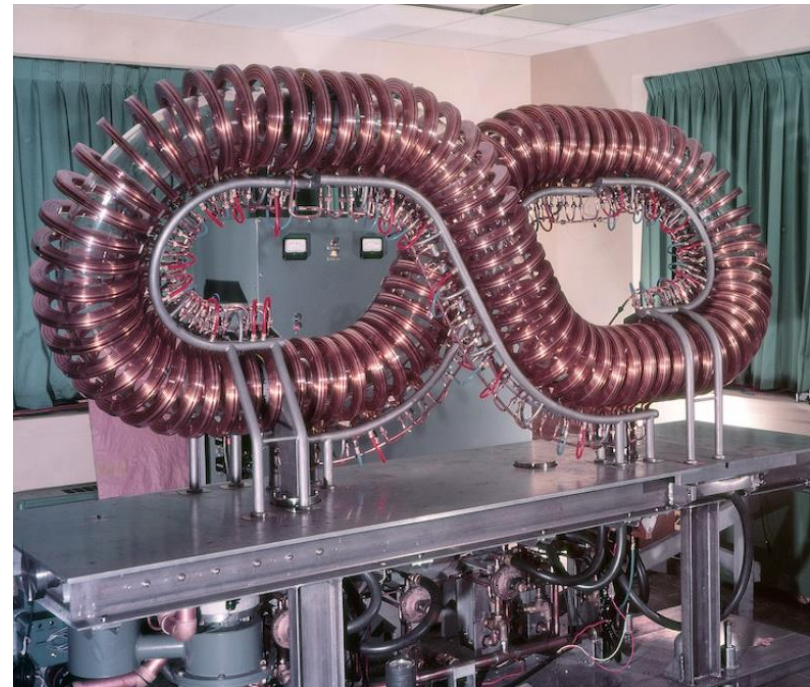
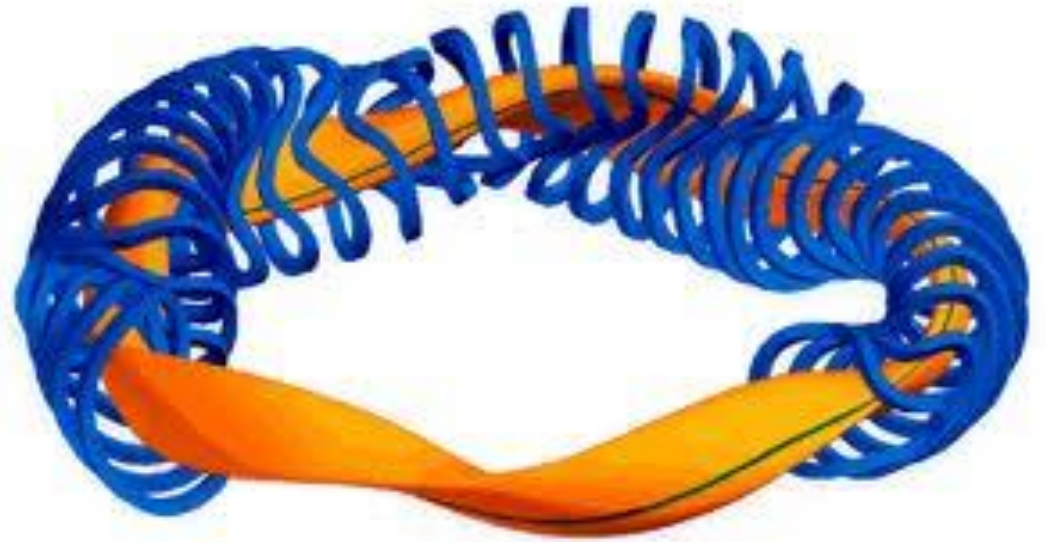


Plasma control via machine learning



Credit: DeepMind & SPC/EPFL

Steady state operation and the resurgence of stellarators



Progress in magnetic confinement fusion

